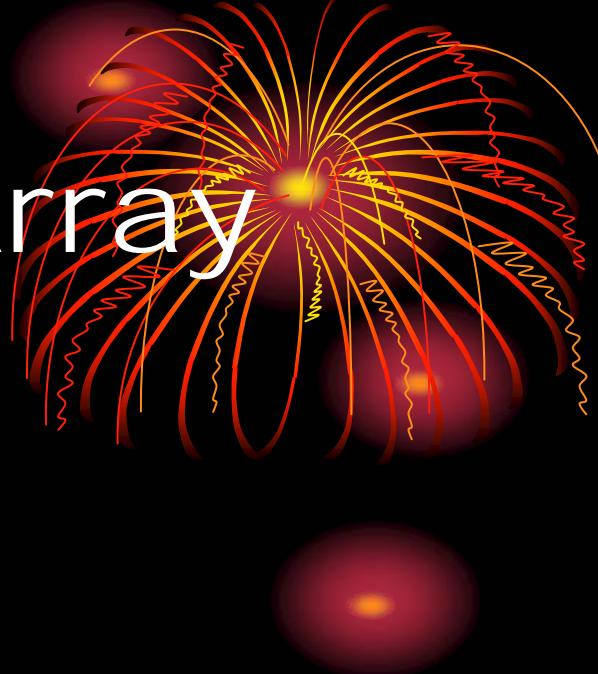


# Extragalactic Science with the Allen Telescope Array

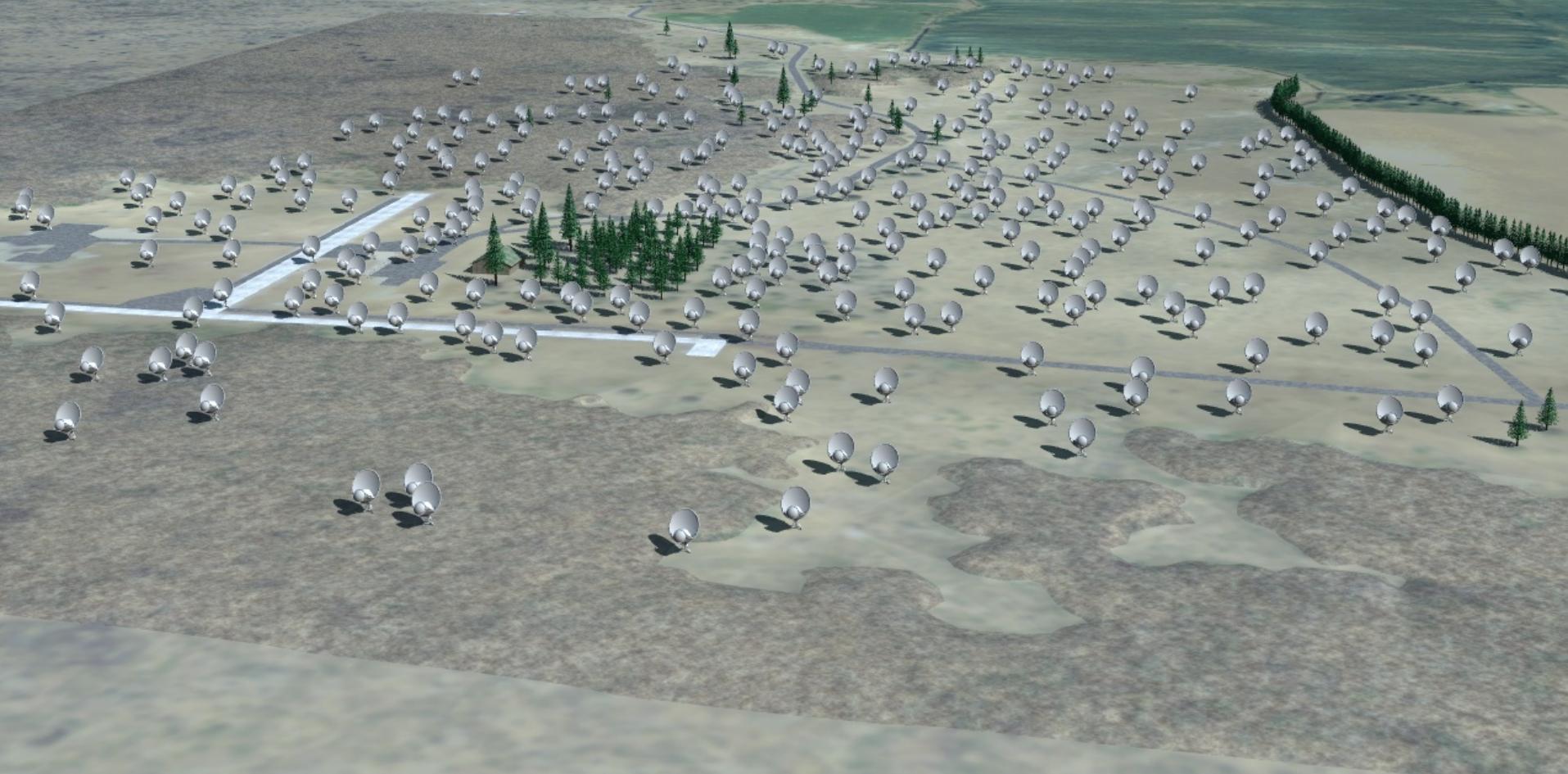
Geoffrey C. Bower  
UC Berkeley

# Allen Telescope Array



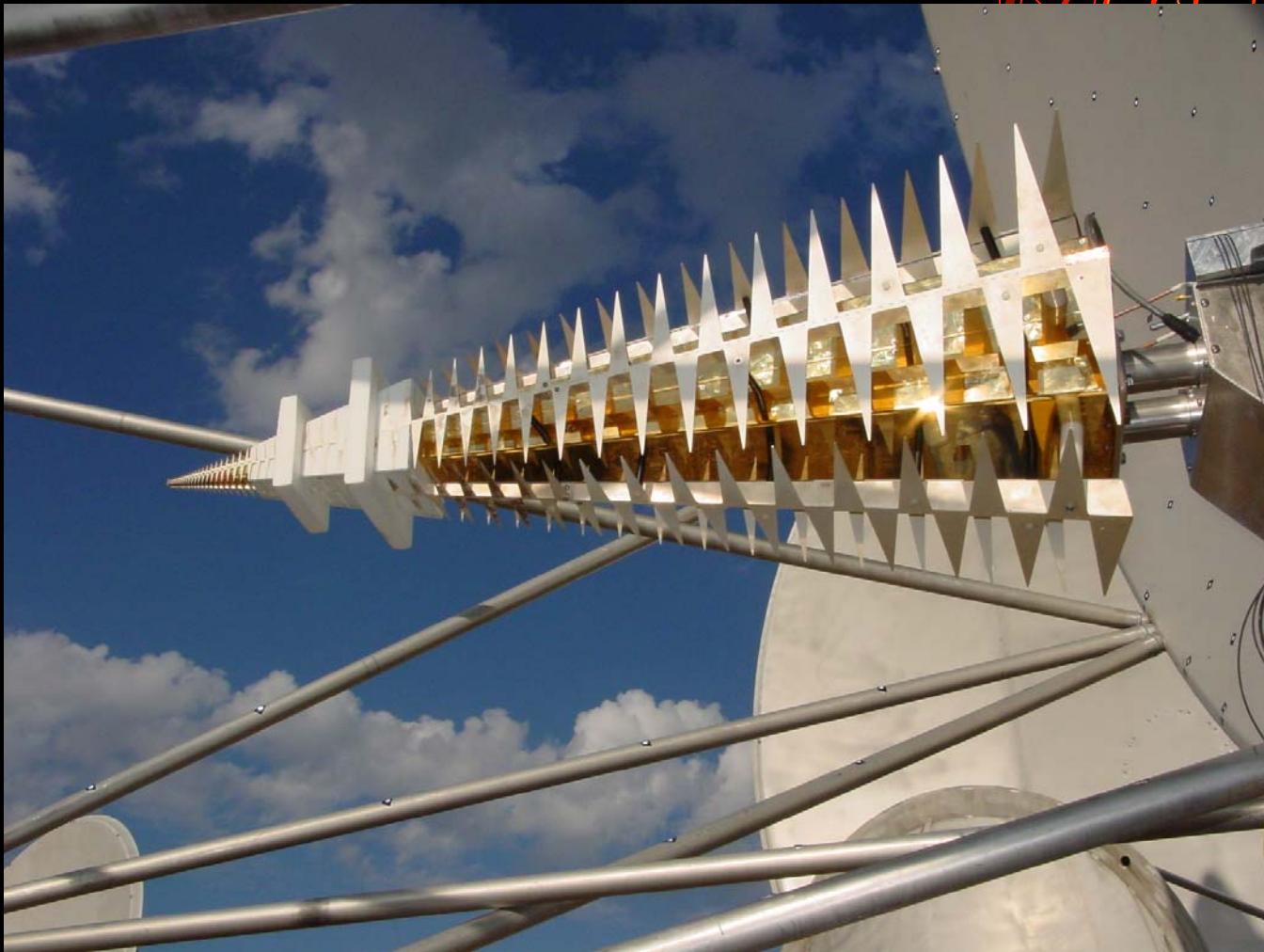
- Large N design
  - 350 x 6.1m antennas
  - Sensitivity of the VLA
- Continuous frequency coverage
  - 0.3 to 11.2 GHz
- Wide field of view
  - 3 degrees at 1 GHz
  - Excellent survey instrument
- Simultaneous observing with multiple backends
  - Correlator at 2 frequencies
  - 16 Phased array beams
- Joint project of UC Berkeley/SETI Inst.
- Prototype for US SKA proposal

# Hat Creek, CA





# Ultrawideband Log Periodic Feed



3 Now

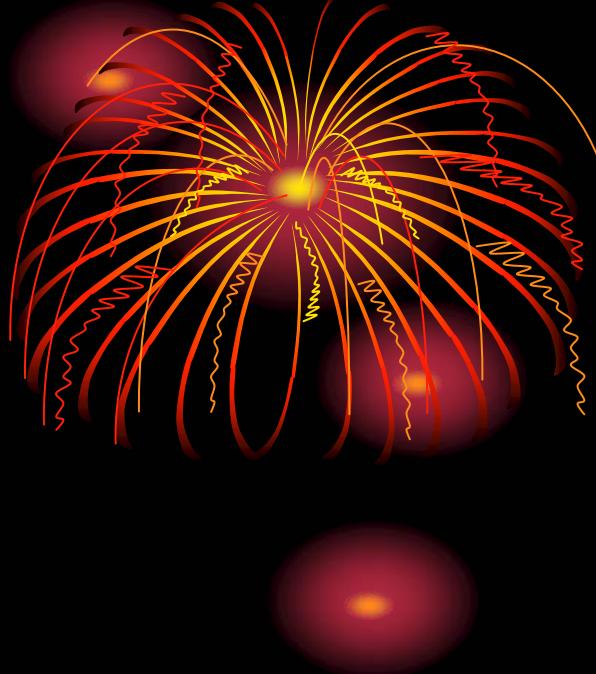
32 under construction, due end 2004

350 by 2007



# ATA Science

- SETI
- Deuterium detection
- Galactic Spectroscopy
- Pulsars
- Surveys, surveys, surveys



# Extragalactic Science

- Dark galaxies
- Radio transients
- Deep, high frequency AGN surveys
- High redshift atomic and molecular absorption



# The HI SDSS Complement

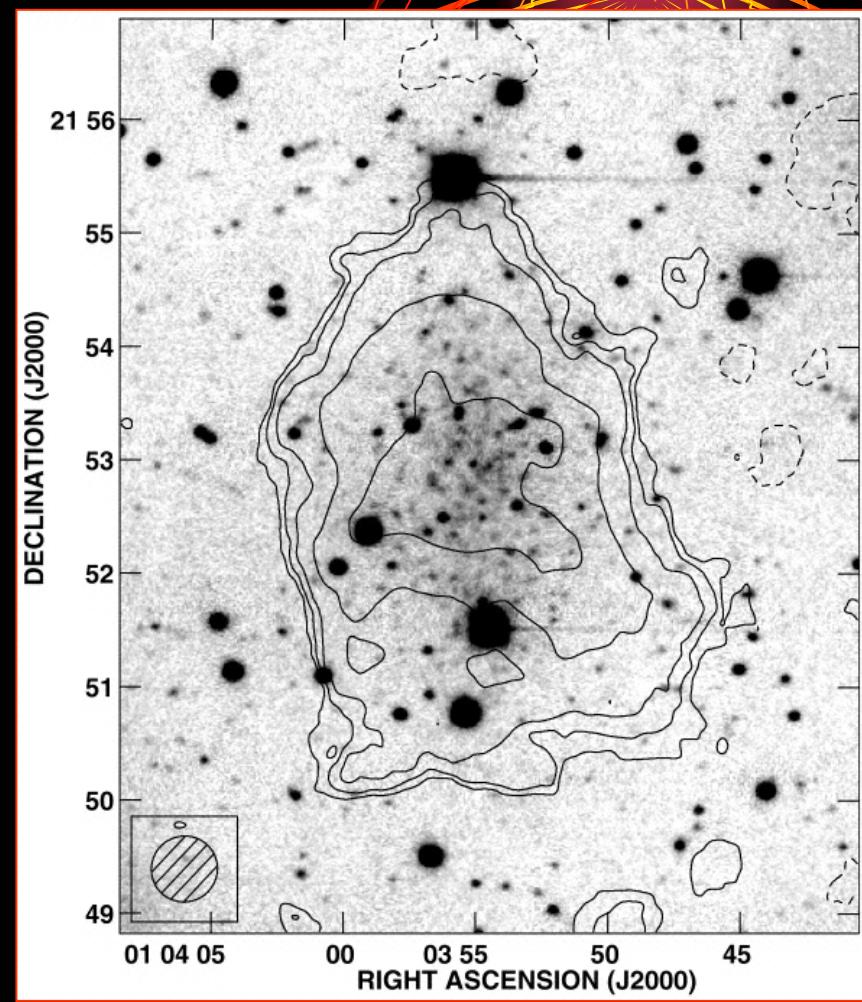
Dark Galaxies to redshift  
0.2



# Dark Galaxies

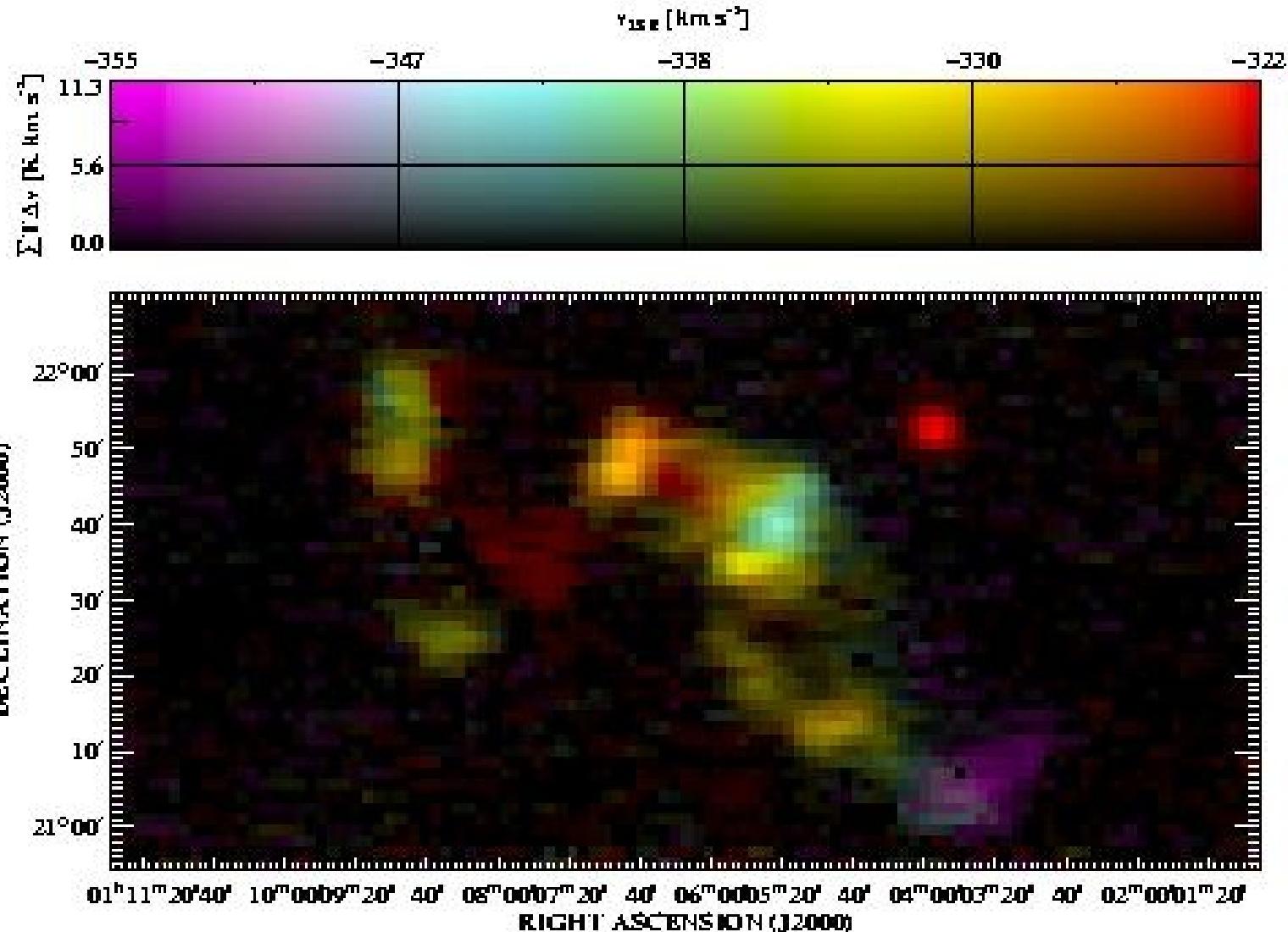


LGS 3



Young & Lo 1997

# Are there dark galaxies?



$2 \times 10^7$  Solar Masses

Robishaw et al 2002

# Consider some ATA sensitivities for HI

Assume:  $\Delta V = 214 \text{ km s}^{-1}$

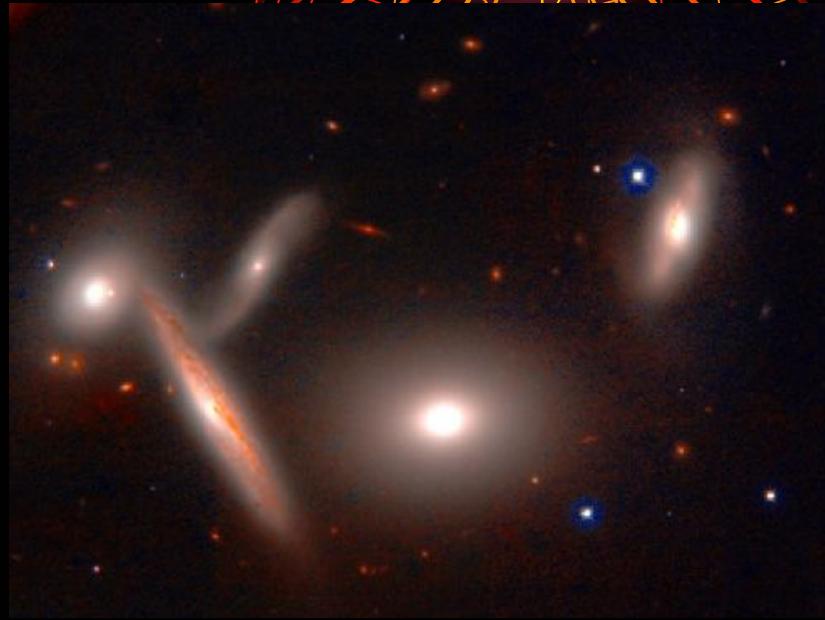
$\Theta = 76''$

4 hour integration

Sensitivity:

70  $\mu\text{Jy}/\text{beam (rms)}$

10 mK



HI Mass Sensitivity: (5 sigma - 2 pols)

$1.9 \times 10^9 M_{\odot}$  @  $500 h_{70} \text{ Mpc}$  ( $cz = 34,000 \text{ km s}^{-1}$ )

can detect  $L_*$  to  $z \sim 0.16$  over entire northern sky  
well matched to Sloan DSS – can be part of

HI survey

# HI SDSS Complement Survey



- *HI Mass Sensitivity:* (5 sigma)
  - $1.9 \times 10^9 M_{\odot}$  @ 500  $h_{70}$  Mpc
  - L\* to z ~ 0.16 over entire northern sky
- rotation curves for all spatially resolved galaxies
- Ultimate zone of avoidance survey
- Structure of the universe at low z
- Detection of dark galaxies
- Confusion-limited AGN survey

# Transient Science with the ATA

- Targeted Monitoring
  - Exploits multiple beams
  - Pulsars
  - Gamma-ray burst afterglows
  - Black holes
  - Supernovae
  - Intraday variability
  - *Your Favorite Object Here*
- Blind Surveys
  - Exploit wide FOV
  - Low-Hanging Fruit
  - Orphan GRB afterglows
  - New Supernovae
  - Tidally disrupted stars around Massive BHs
  - The Twinkling Radio Sky
  - *Something New!*

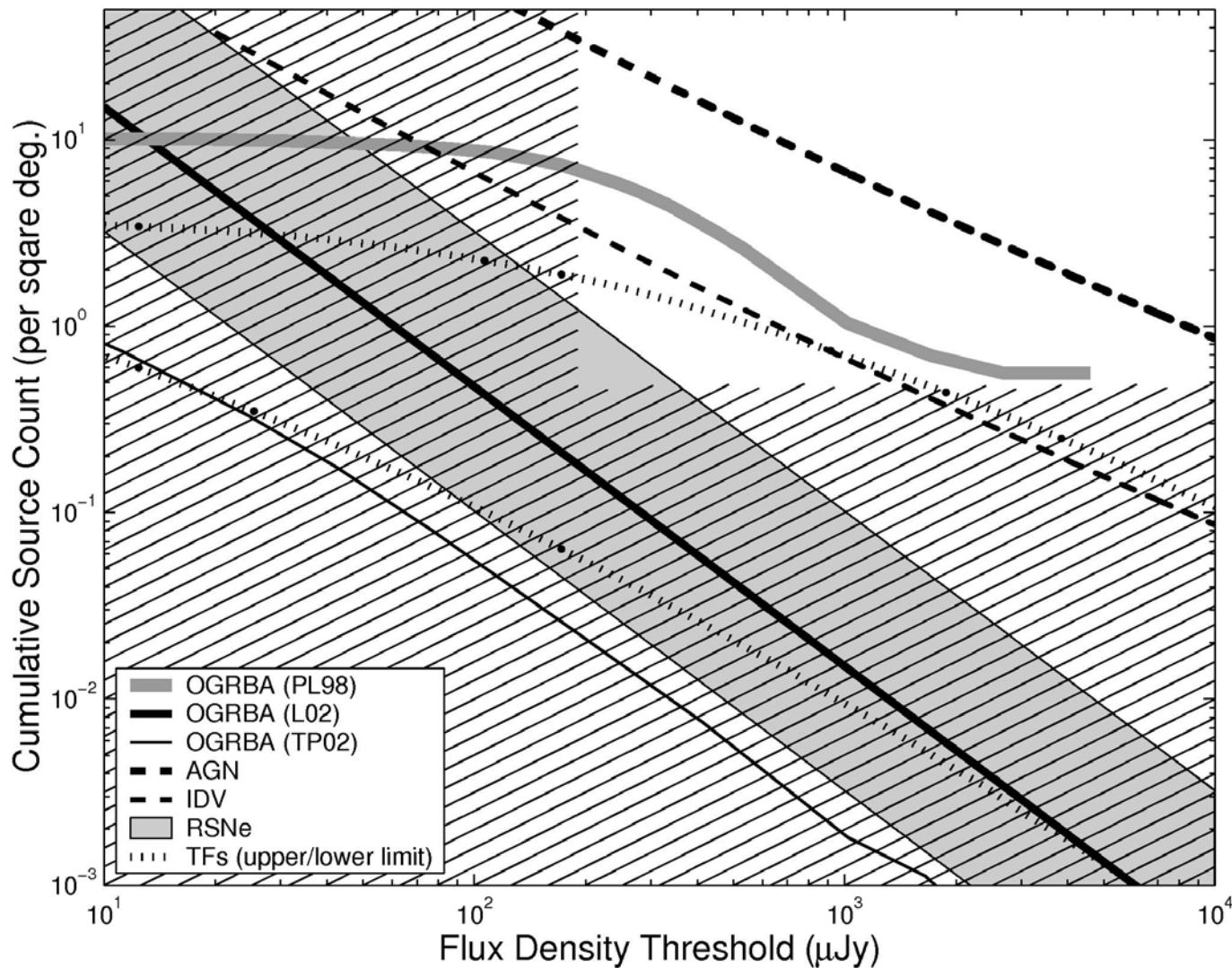


# Orphan Gamma Ray Burst Afterglows



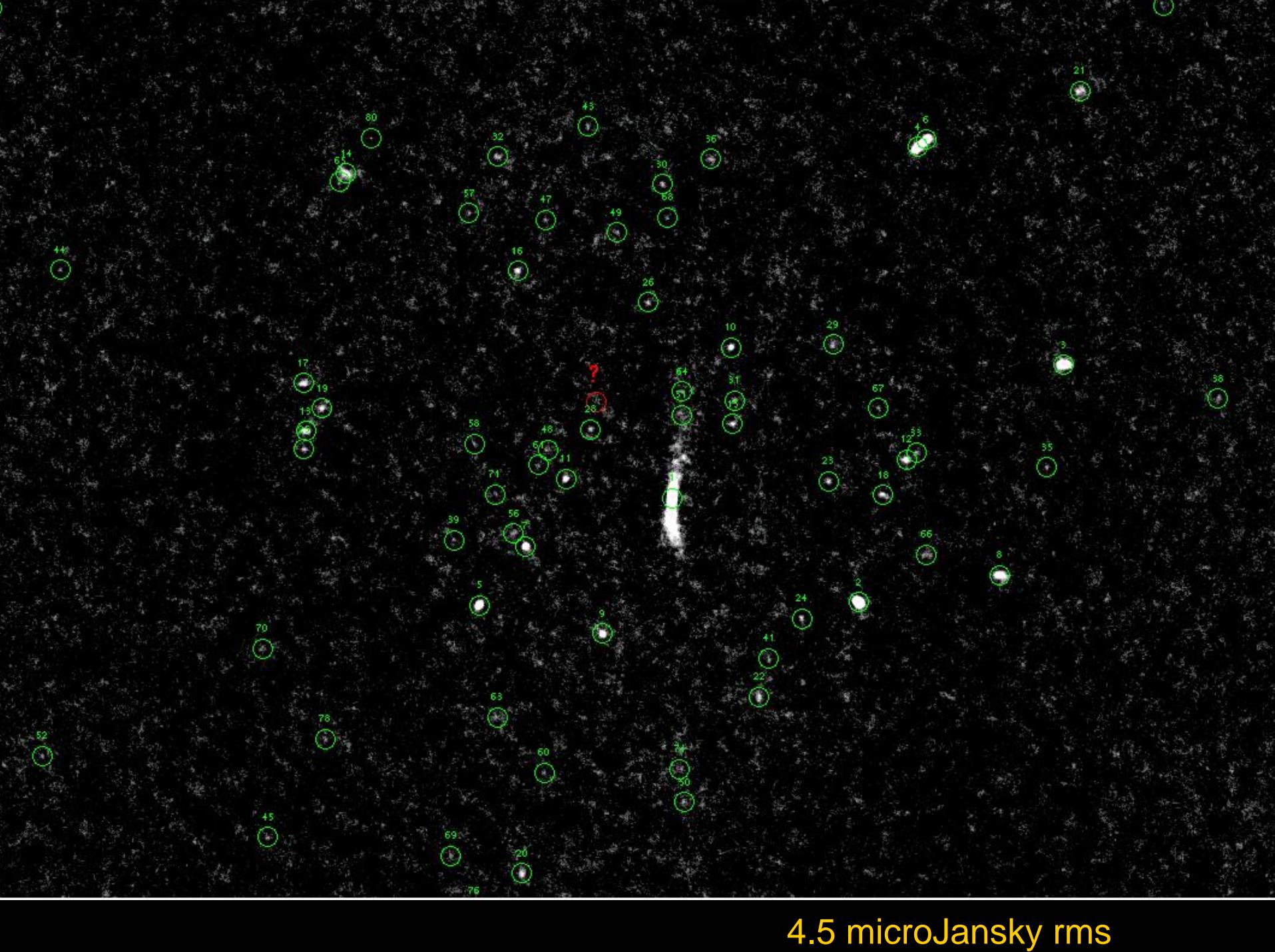
- $\gamma$ - ray quiet, radio loud bursts
  - Not yet discovered
- Search tests basic GRB model
- Find the first stars and black holes ( $z \sim 10-30$ )
- Identify the foreground for Square Kilometer Array

# Radio Transient Source Counts



VLA  
200  
hours



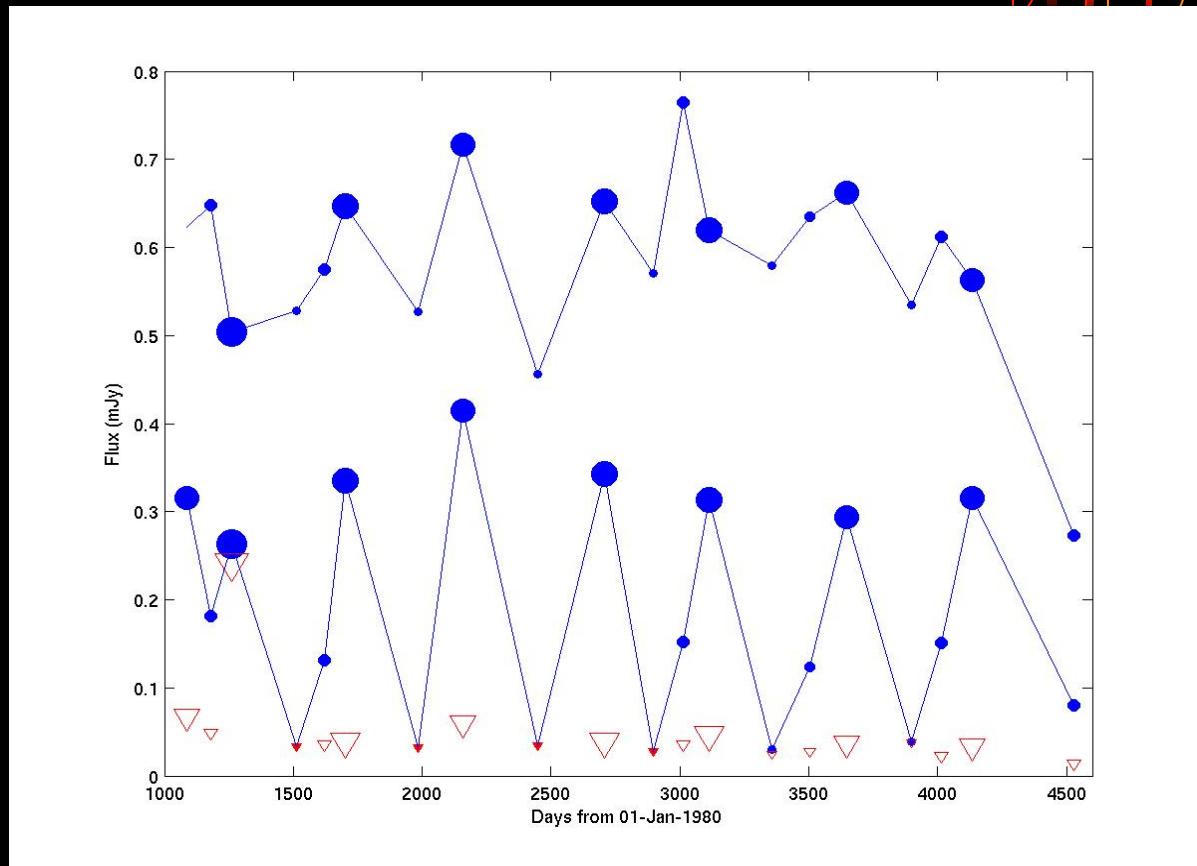


# VLA Archive Transient Survey

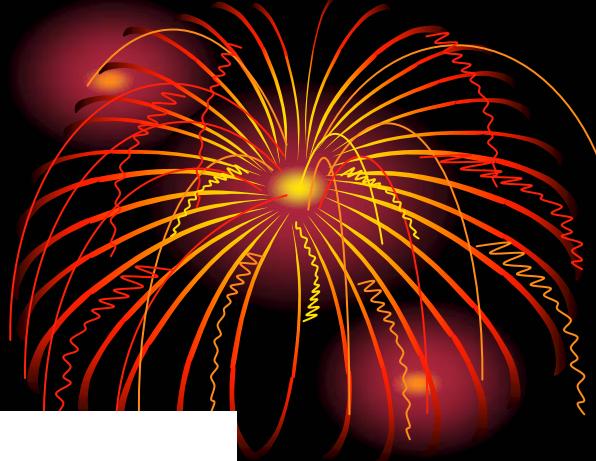
- 10 years
- 20 epochs
- 5 GHz
- 20 microJansky rms/epoch
- 80 sources
  - 20 suitable for transient study



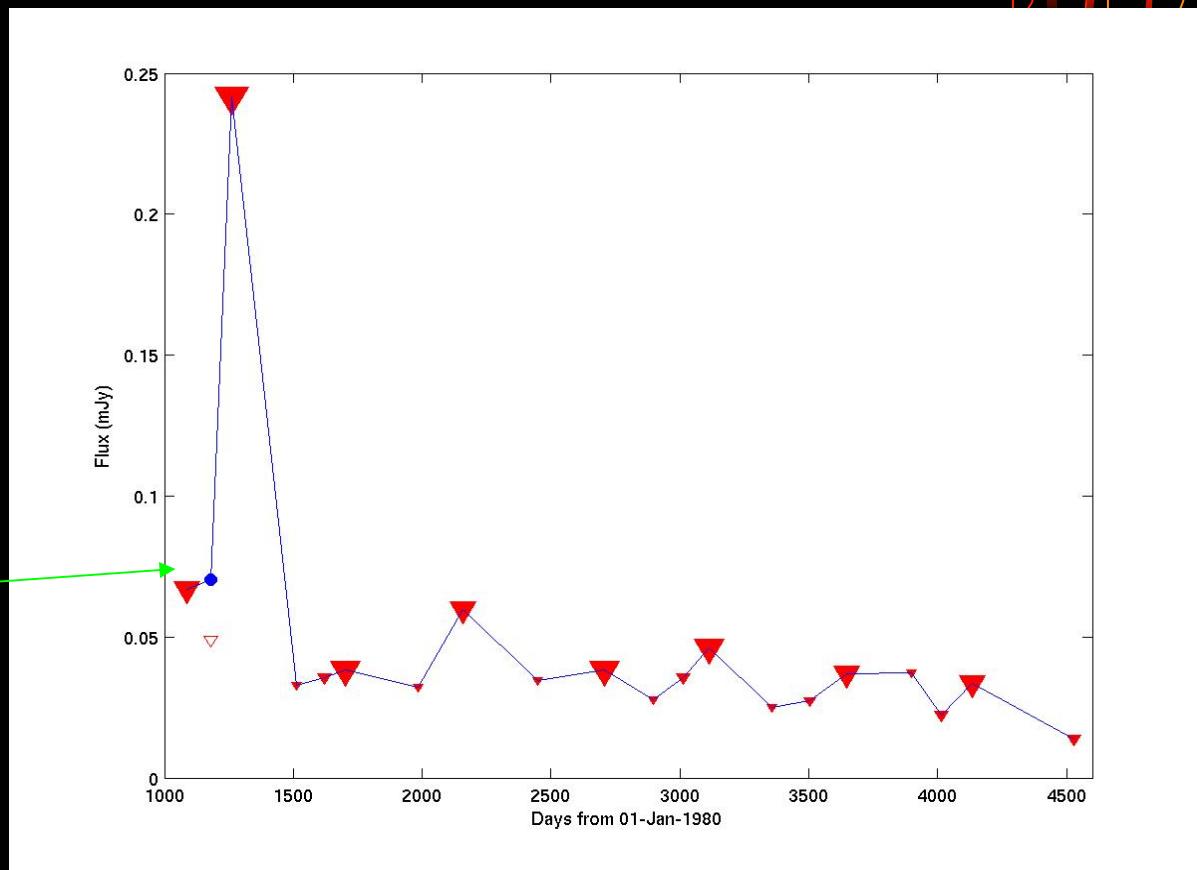
# Sample Light Curves



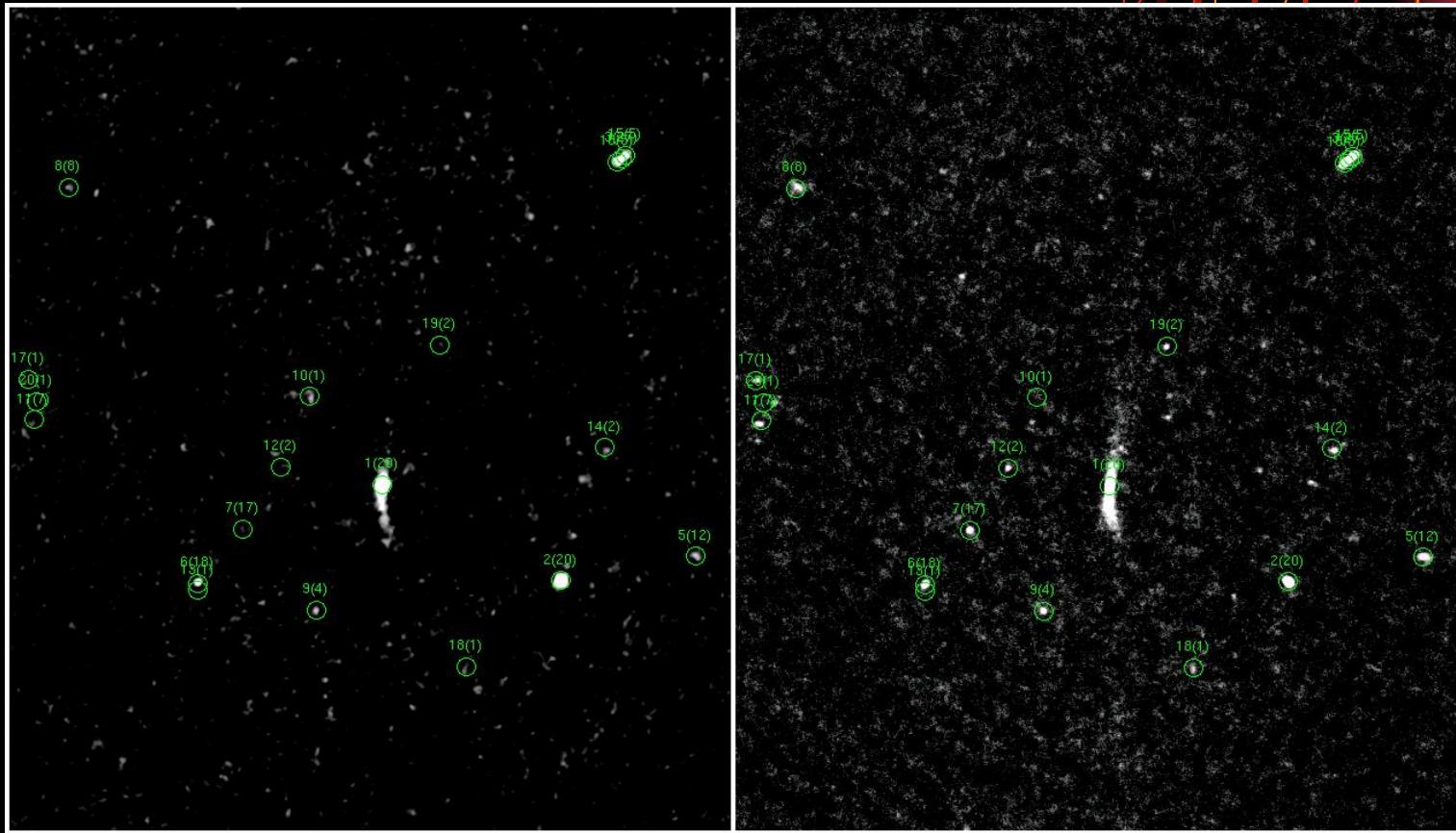
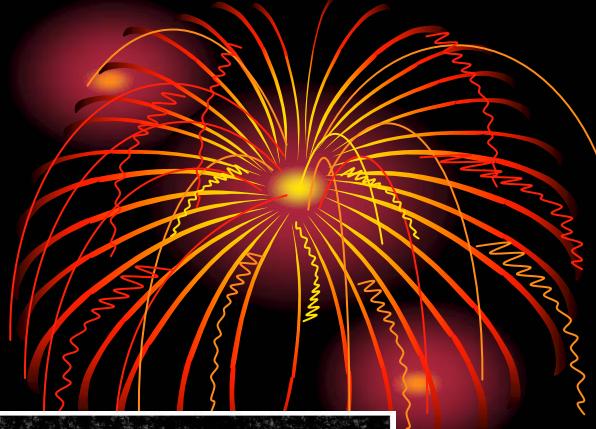
# Transient?



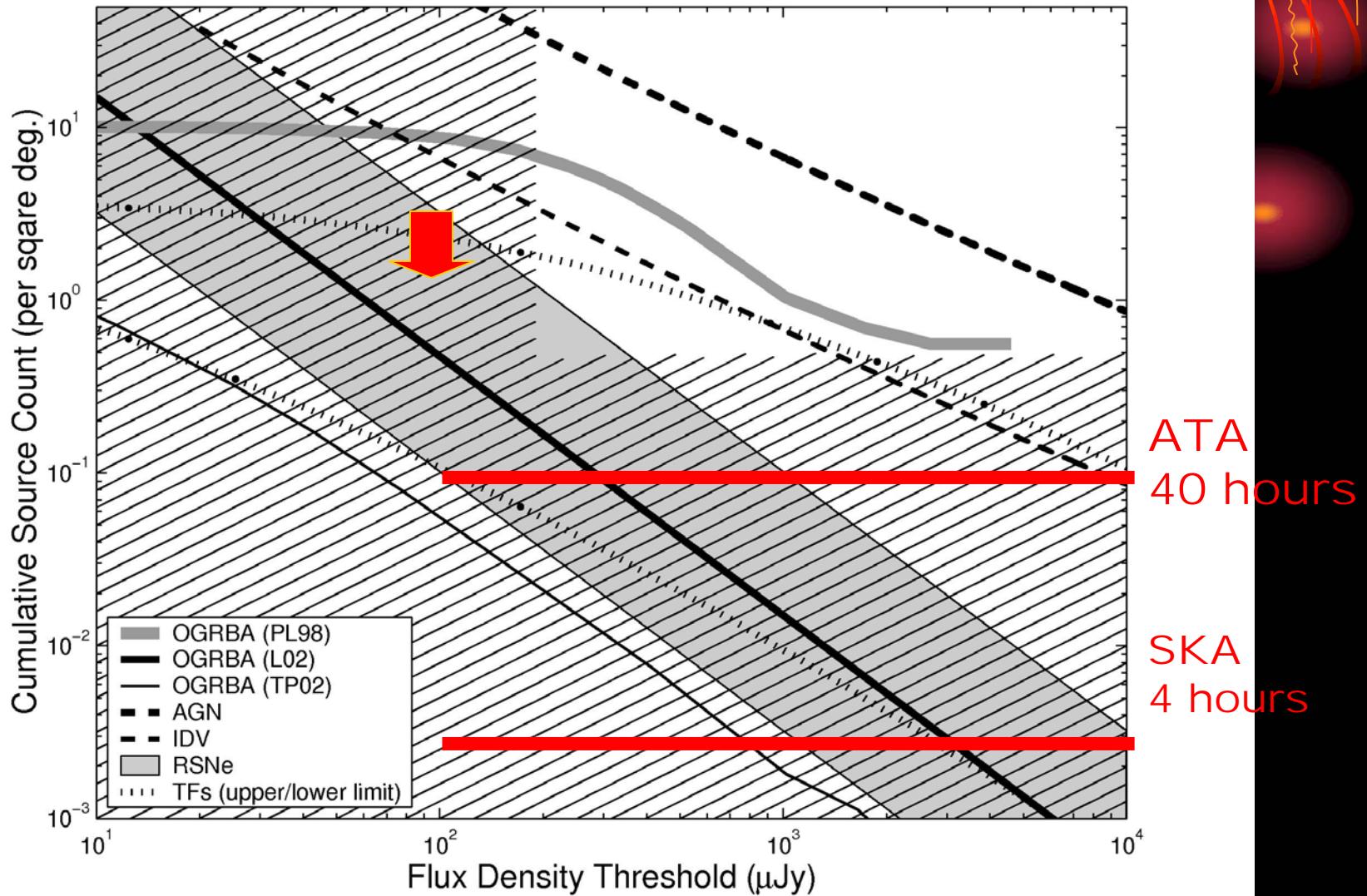
5 Sigma  
Detection



# Transient?



# Radio Transient Source Counts

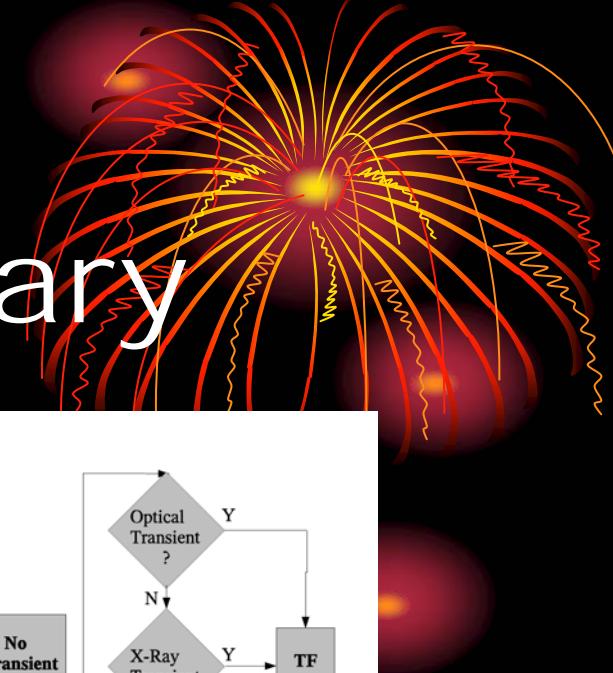
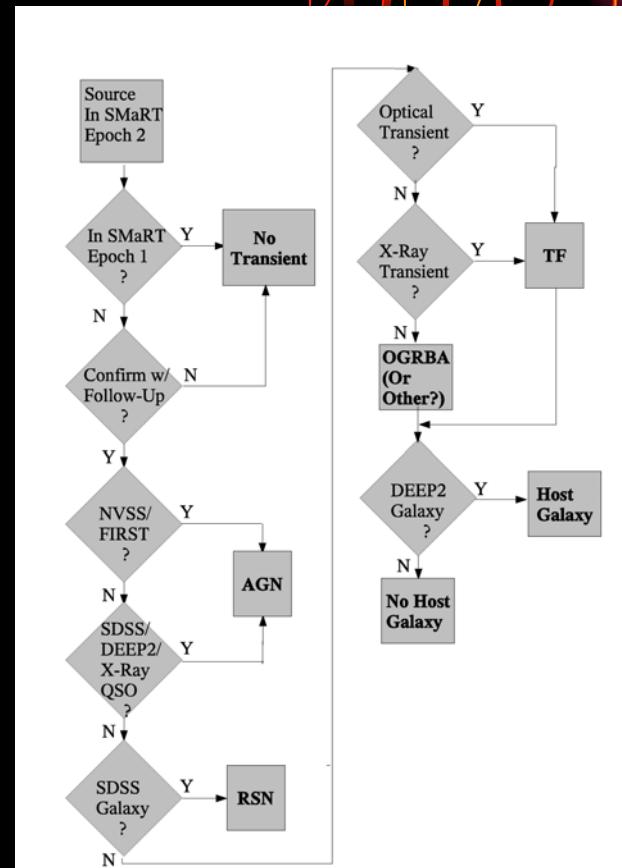


ATA  
40 hours

SKA  
4 hours

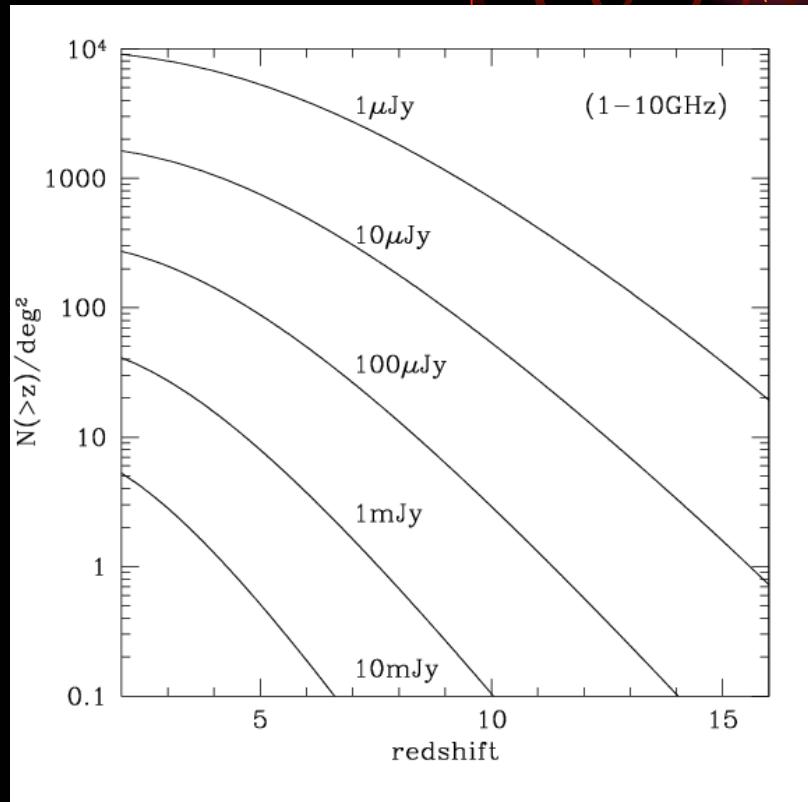
# Multi-wavelength Approach Necessary

- Many types of sources are transient
- Multi-lambda necessary to determine type

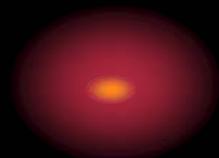


# High Redshift AGN Surveys

- ~ $10^4$  FIRST sources (~1%) at  $z > 8$
- ATA can conduct deep & wide area survey
  - 6 GHz,  $10^4$  square degree in the same time



# ATA will be a powerful instrument for extragalactic surveys



- HI SDSS complement
  - Dark galaxies
  - Structure of the local universe
- Transients
  - Terra incognita
  - Potentially very high redshift science