

Black hole candidates in the nearby universe

Speaker:

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COST

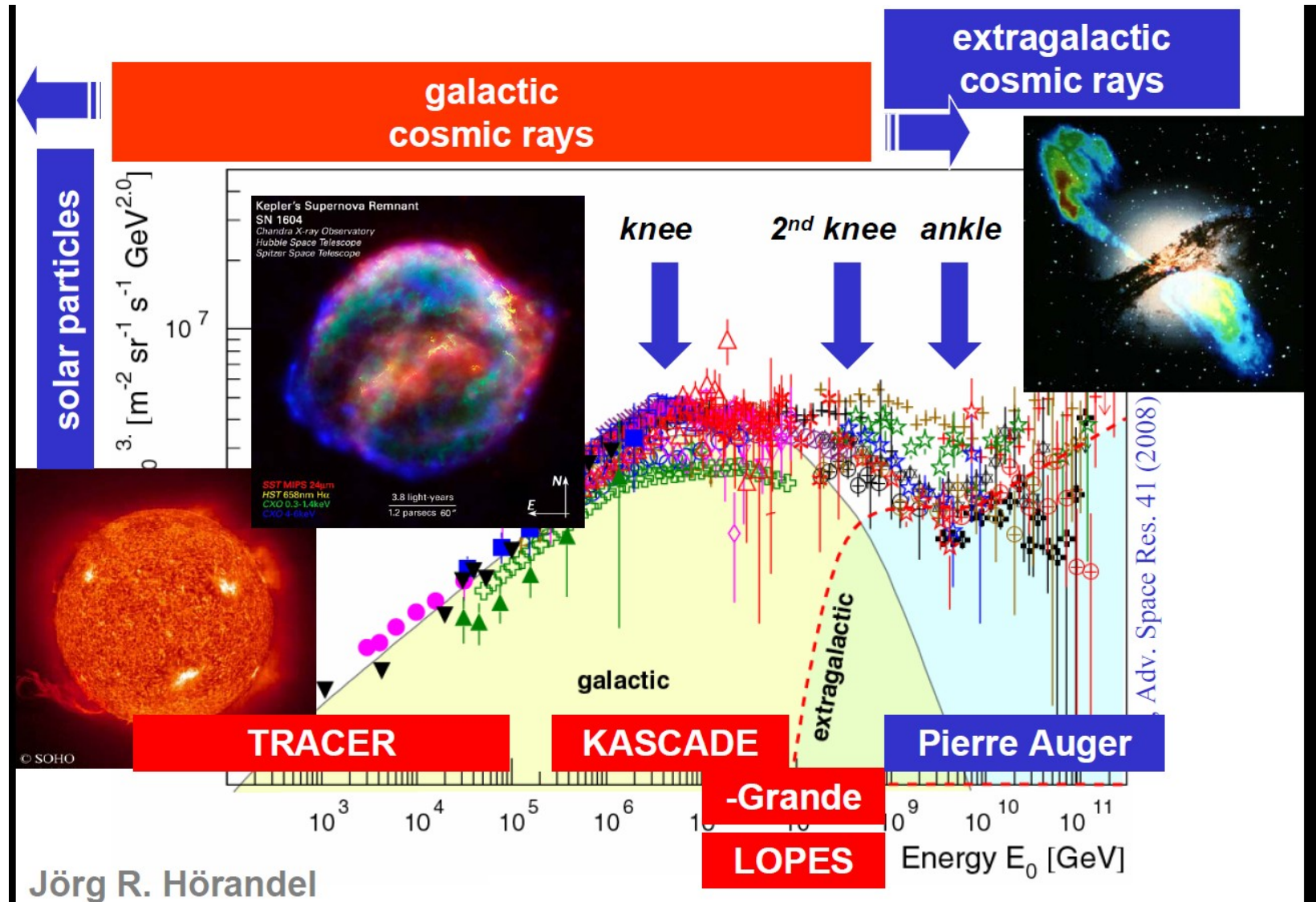
WG 4 Supermassive
Black Holes



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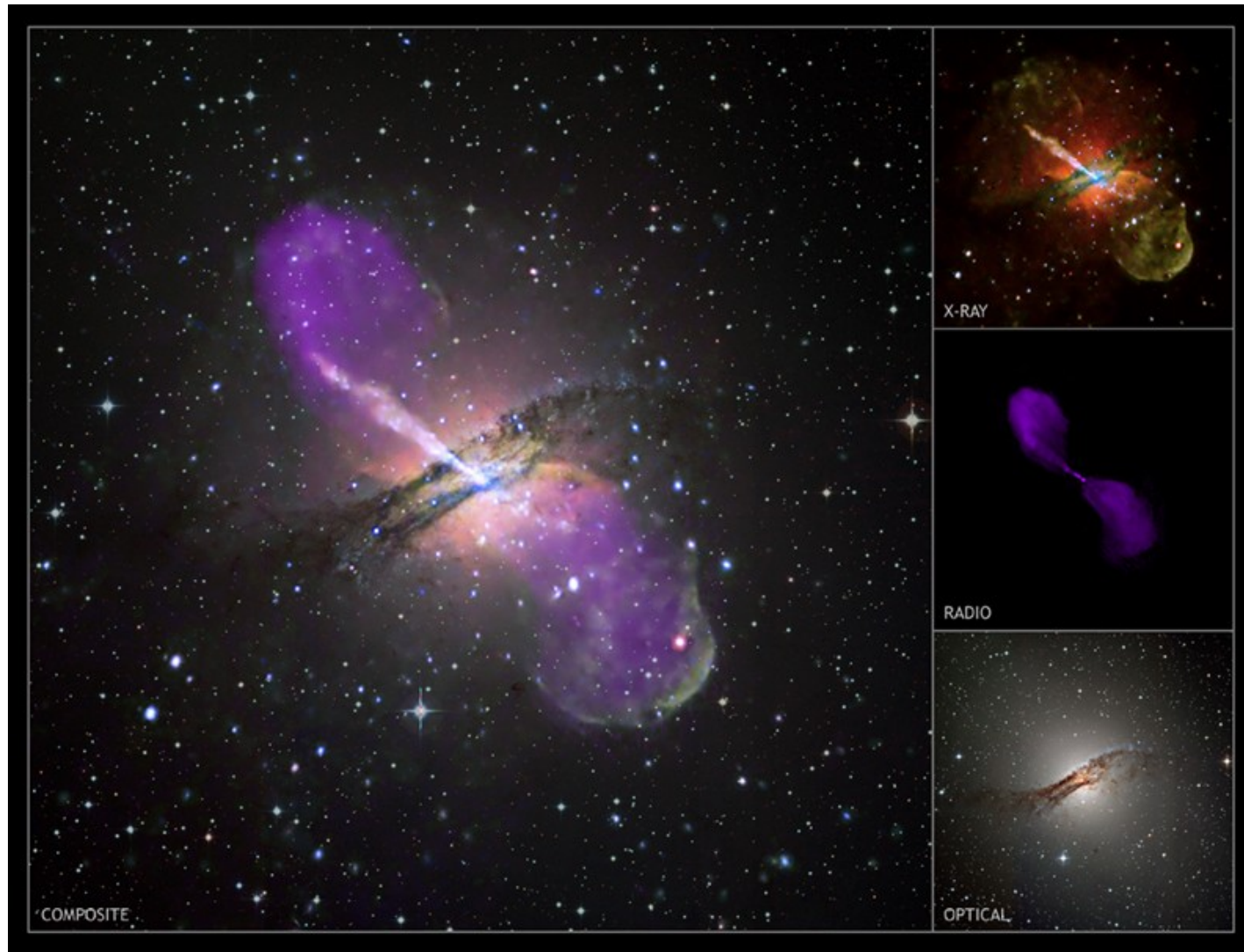
- From Cosmic Rays to Black Holes
 - Massive Black Holes Catalog
 - Characteristics of the catalog
 - Monte-Carlo Distribution Simulation
 - Conclusions&Discussions
-

- UHECR – Sources



Jörg R. Hörandel

- UHECR – Sources



Active Galaxy Centaurus A

Credit: X-ray - NASA, CXC, R.Kraft (CfA), et al.;

Radio - NSF, VLA, M.Hardcastle (U Hertfordshire) et al.;

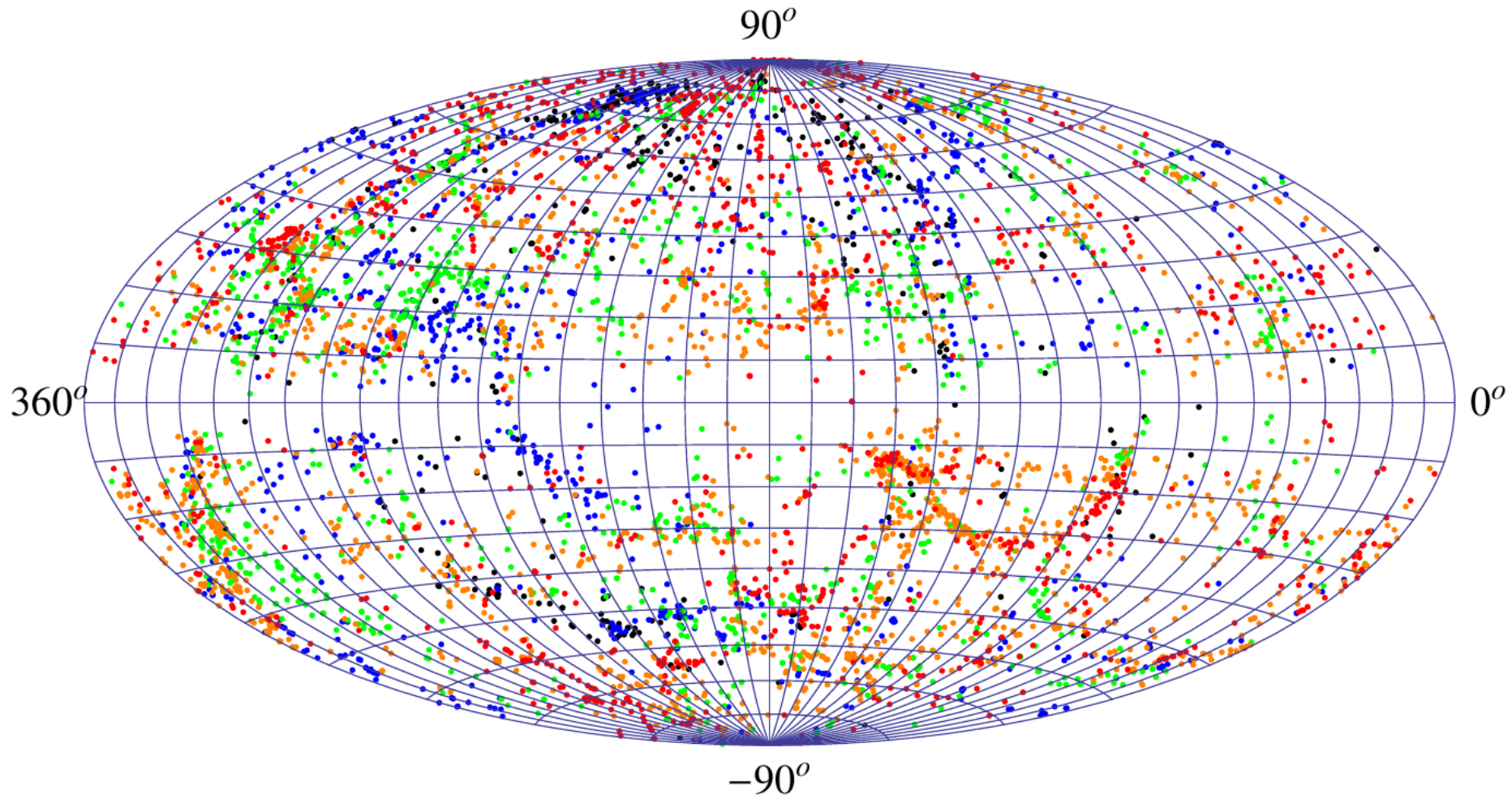
Optical - ESO, M.Rejkuba (ESO-Garching) et al.

- Constructing the catalog

- Sample at 2 micron (Two Micron All Sky Survey - 2MASS, Skrutskie et al., 2006): all galaxies dominated by an **old stellar population** emit strongly at **2micron** (van der Wel et al. 2006), and the spheroidal component of the **older stellar population** correlates well with **supermassive black holes** (e.g., Faber et al. 1997, Wang & Biermann 1998, Haering & Rix 2004)

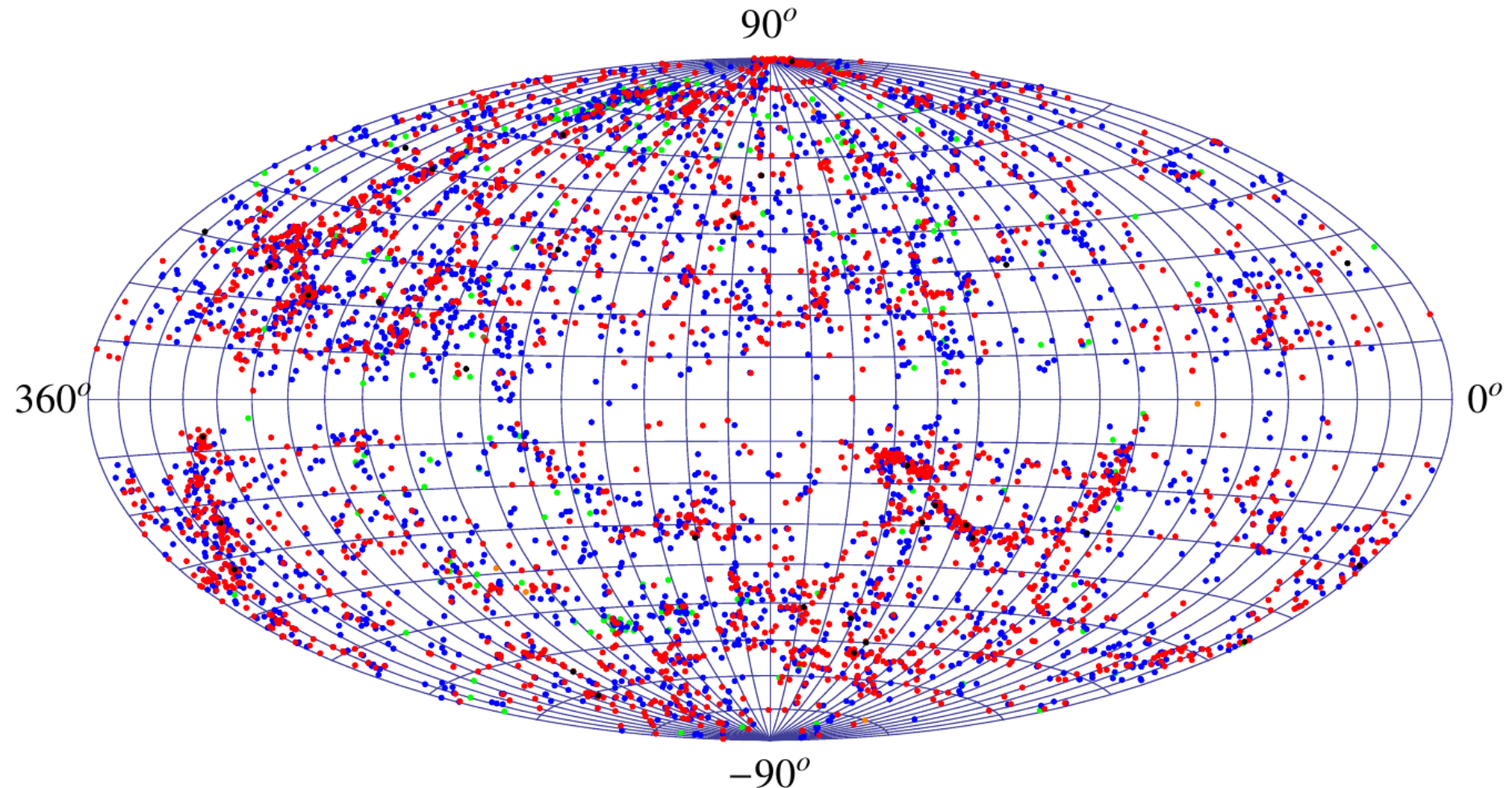
• Using Hubble type \longrightarrow • Scaling Relation \longrightarrow • Catalog

- Sky plot color coded in Redshift



Aitoff projection in galactic coordinates of 5,895 NED candidate sources in the case of a complete sub sample. The choice was made from a complete sample of 10284 candidate brighter than 0.03Jy and selected at $z < 0.025$ and 2 micron. The color code is Black, Blue, Green, Orange, Red corresponding to redshifts between 0, 0.005, 0.01, 0.015, 0.02, 0.025

- Sky plot color coded in Black Holes Mass



Aitoff projection in galactic coordinates of 5,895 NED candidate sources. The color code is Brown, Orange, Green, Blue, Red, Black corresponding to black hole masses between 10^5 Msol , 10^6 Msol , 10^7 Msol , 10^8 Msol , 10^9 Msol .

• Example of black holes list

| Name | l Deg | b Deg | z | Morphological type | Distance Mpc | Estimated M_{BH} $10^8 M_{\odot}$ | B-V | FIR/Radio ratio |
|----------------|----------|----------|----------|-----------------------|-----------------|--|-------|--------------------|
| NGC 5332 | 0.16 | 72.678 | 0.02241 | S0- | 92.0 | 3.75 ± 1.88 | 0.94 | |
| NGC 6500 | 43.763 | 20.233 | 0.01001 | SAab LINER | 41.1 | 1.41 ± 0.704 | | 3.6 ± 0.1 |
| NGC 6849 | 0.329 | -30.818 | 0.02014 | SB0- | 82.7 | 4.49 ± 2.24 | 0.8 | |
| NGC 5311 | 83.759 | 72.474 | 0.009 | S0/a | 36.9 | 1.04 ± 0.521 | | 9.4 ± 0.9 |
| NGC 5845 | 0.338 | 48.904 | 0.00483 | E | 24.1 | 0.50 ± 0.25 | 0.97 | < 0.17 |
| NGC 5850 | 0.516 | 48.636 | 0.00852 | SB(r)b | 35.0 | 2.68 ± 1.34 | 0.72 | < 0.819 |
| NGC 7469 | 83.099 | -45.467 | 0.01631 | (R')SAB(rs)a Sy1.2 | 67.0 | 4.96 ± 2.48 | 0.55 | 384.9 ± 57.1 |
| NGC 5846 | 0.426 | 48.797 | 0.00571 | E0-1;LINER HII | 29.1 | 5.37 ± 2.69 | 0.96 | > 181.8 |
| ESO 338- G 009 | 0.445 | -23.401 | 0.01857 | Sa-b | 76.2 | 1.16 ± 0.582 | | |
| NGC 5838 | 0.729 | 49.319 | 0.00453 | SA0-LINER | 18.6 | 1.23 ± 0.614 | 0.94 | < 0.73 |
| MESSIER 094 | 123.363 | 76.007 | 0.001027 | (R)SA(r)ab;Sy2 LINER | 4.6 | 0.752 ± 0.376 | 0.72 | 601.1 ± 75.06 |
| UGC 00542 | 123.457 | -33.599 | 0.015044 | Sb | 61.7 | 1.32 ± 0.661 | 13.18 | < 0.22 |
| NGC 4648 | 123.818 | 42.691 | 0.004717 | E3 | 19.3 | 0.413 ± 0.207 | 0.89 | |
| UGC 00555 | 123.823 | -33.999 | 0.022699 | S0/a | 93.2 | 2.67 ± 1.34 | | |
| UGC 01039 | 123.896 | 22.220 | 0.017616 | Sab | 72.3 | 1.41 ± 0.706 | | < 0.43 |
| UGC 00567 | 123.897 | -31.115 | 0.020291 | S0 | 83.3 | 1.69 ± 0.844 | | |
| UGC 00670 | 123.913 | 12.758 | 0.015948 | SBb? | 65.4 | 2.33 ± 1.17 | | < 0.73 |
| NGC 0317A | 124.118 | -19.057 | 0.017656 | S0 | 72.5 | 2.09 ± 1.04 | | < 6.45 |
| UGC 00600 | 124.118 | -14.194 | 0.022726 | SAB(s)b | 93.3 | 2.06 ± 1.03 | | < 0.84 |
| NGC 4589 | 124.235 | 42.901 | 0.006605 | E2 LINER | 20.4 | 1.25 ± 0.627 | 0.93 | 7.1 ± 0.42 |

Example of the data included in the massive black hole catalog, these 20 galaxies illustrate some of the different parameters that are available. We mention here that we used compiled distances as noted before where available and for all distances similar or less than the Virgo cluster the distances are corrected.

- Hubble types

| Type | 5894 selection | 2928 selection | Rejected list |
|--------------|----------------|----------------|---------------|
| E | 783 | 765 | - |
| S0 | 2626 | 1771 | - |
| Sa | 879 | 184 | - |
| Sb | 1052 | 128 | - |
| Sab | 554 | 80 | - |
| Sc | - | - | 1941 |
| Irr | - | - | 14 |
| Sy | - | - | 128 |
| No type | - | - | 867 |
| Sbrst | - | - | 421 |
| Unknown type | - | - | 1019 |

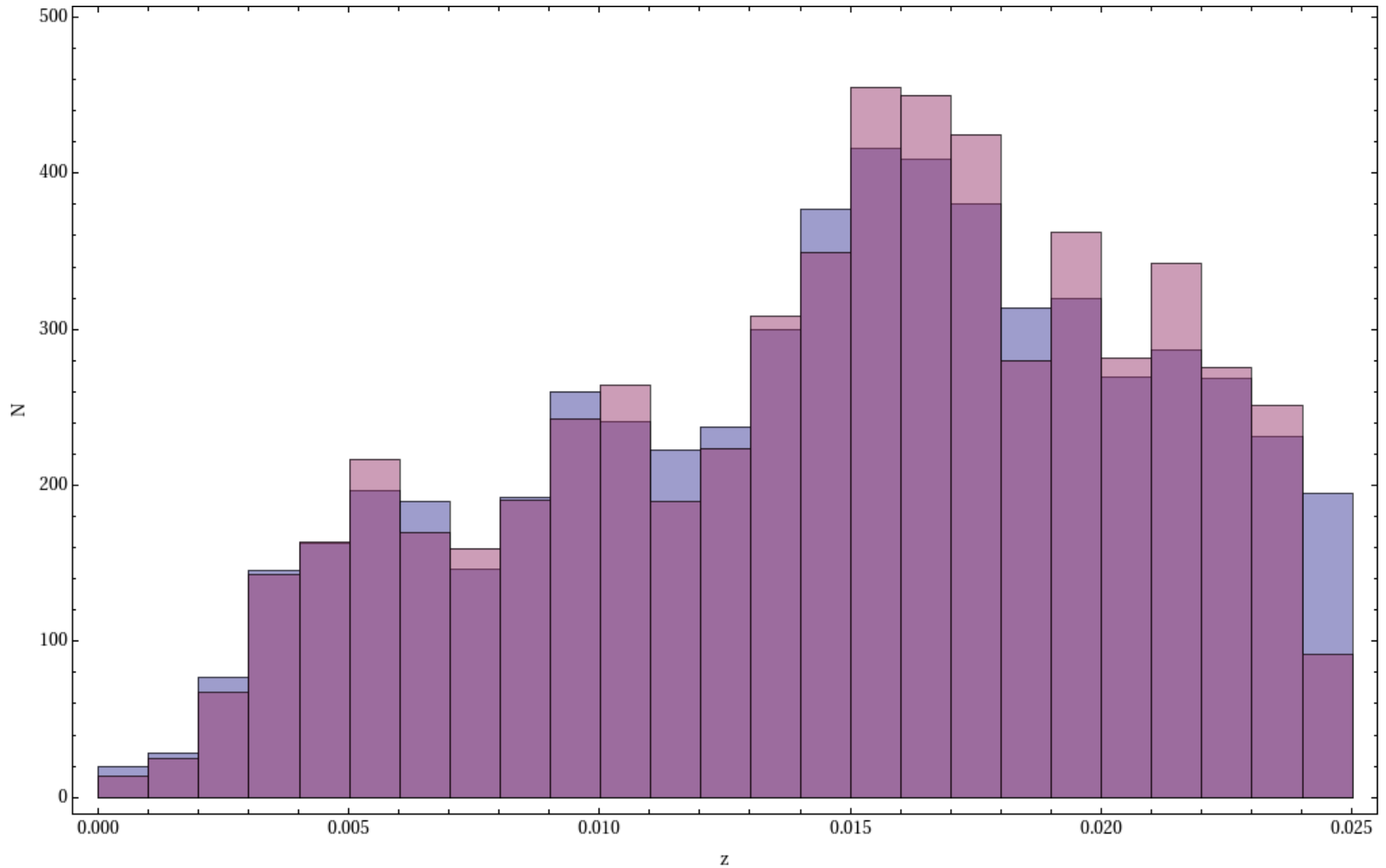
- Application to black hole mass distribution

☺ Importance Sampling

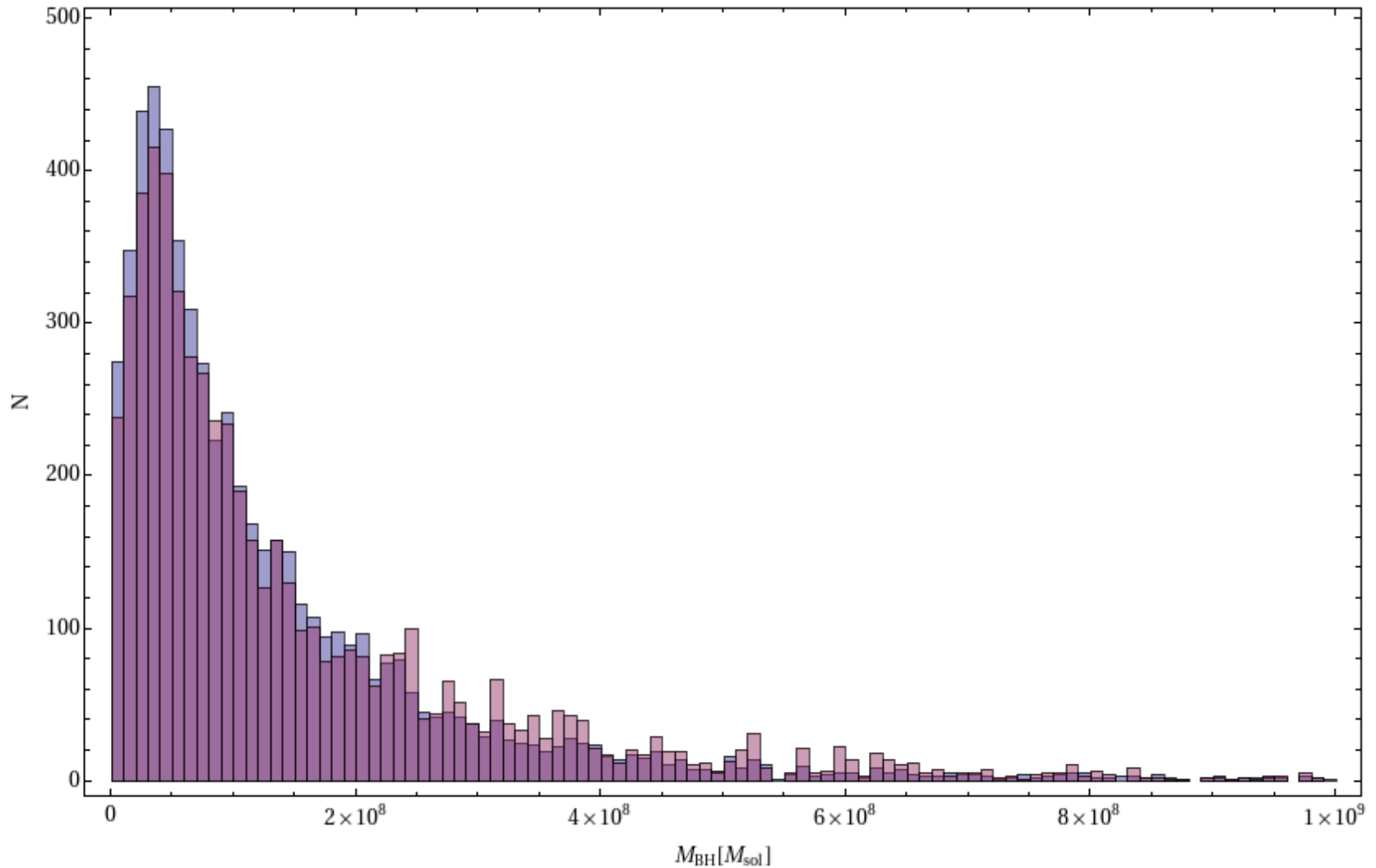
☺ Rejection Sampling

☺ Markov Chain Monte-Carlo Metropolis Hastings

• Markov Chain Monte-Carlo



• Markov Chain Monte-Carlo



- Conclusions&Discussion

- A simple scaling method to give a big catalog of black holes
- Lots of upgrades for the catalog possible after observations in radio, optical and infrared
 - Monte-Carlo multi-dimensional simulations of distribution of mass, redshift, luminosity, etc. of black holes or black holes host galaxies

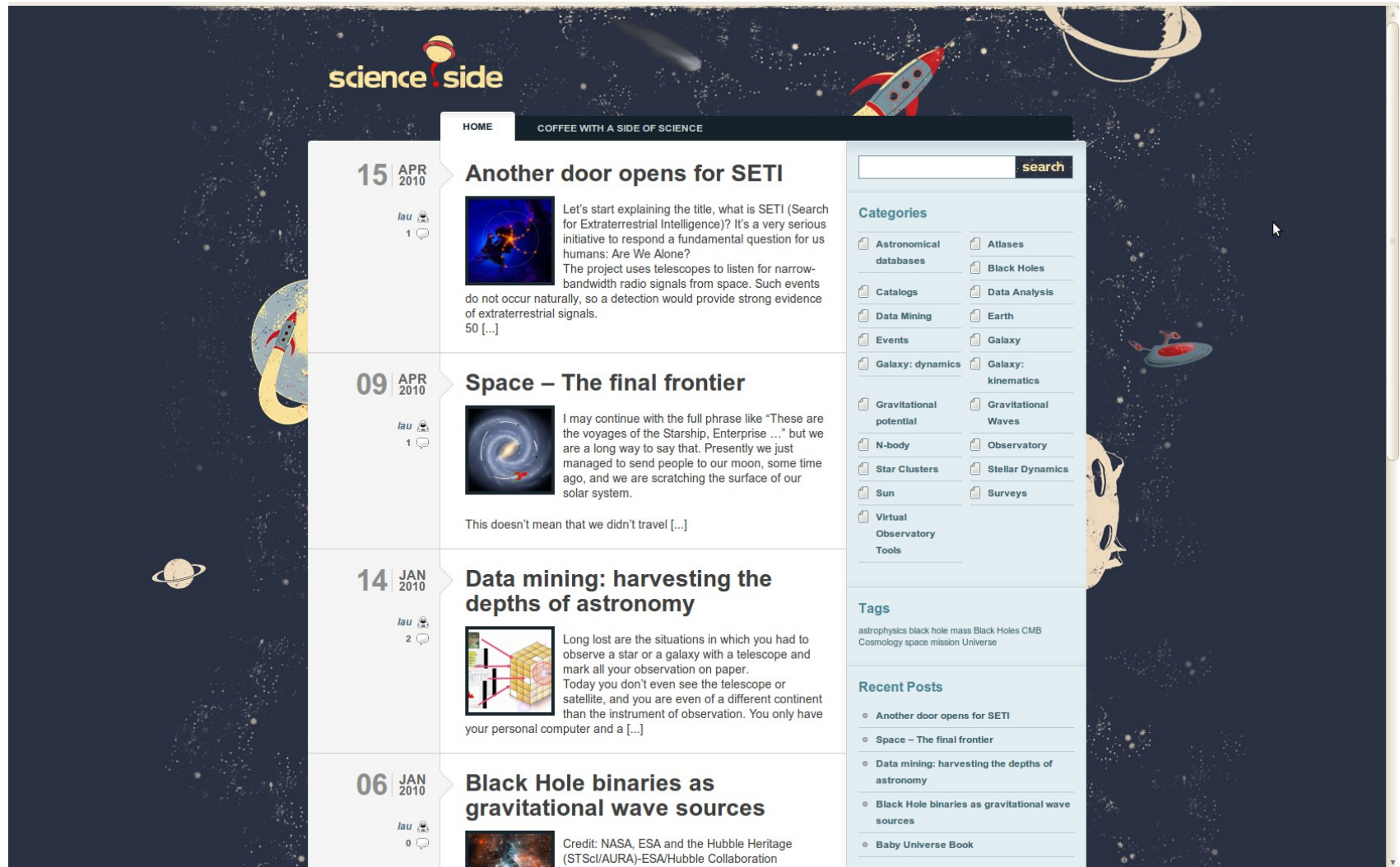
- More info
-

[**http://www.science-side.com/**](http://www.science-side.com/)

[**http://www.spacescience.ro/new1/cosmo/**](http://www.spacescience.ro/new1/cosmo/)


[**http://www.mpifr-bonn.mpg.de/div/theory/**](http://www.mpifr-bonn.mpg.de/div/theory/)

• More info




• More info

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Black Hole binaries as gravitational wave sources



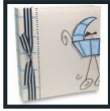
Credit: NASA, ESA and the Hubble Heritage (STScI/AURA)-ESA/Hubble Collaboration

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The main advantage of using the GPU [...]

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Baby Universe Book




Did you know that cosmologists can make a "memory album" of the Universe in its infancy? A so called "Baby Book" which can contain pictures of the Universe when it was as young as 380.000 years old, out of which one can extract useful informations about the birth and evolution of Baby Universe...

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How To Weight A Black Hole



Do you know what a black hole is? Well yes, everybody knows by now what this monster is, an astrophysical object that can absorb everything, even light and from the insight of it there is no way out. But do you know how to measure it's mass? In a recent article by B. Czerny and M. [...]

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
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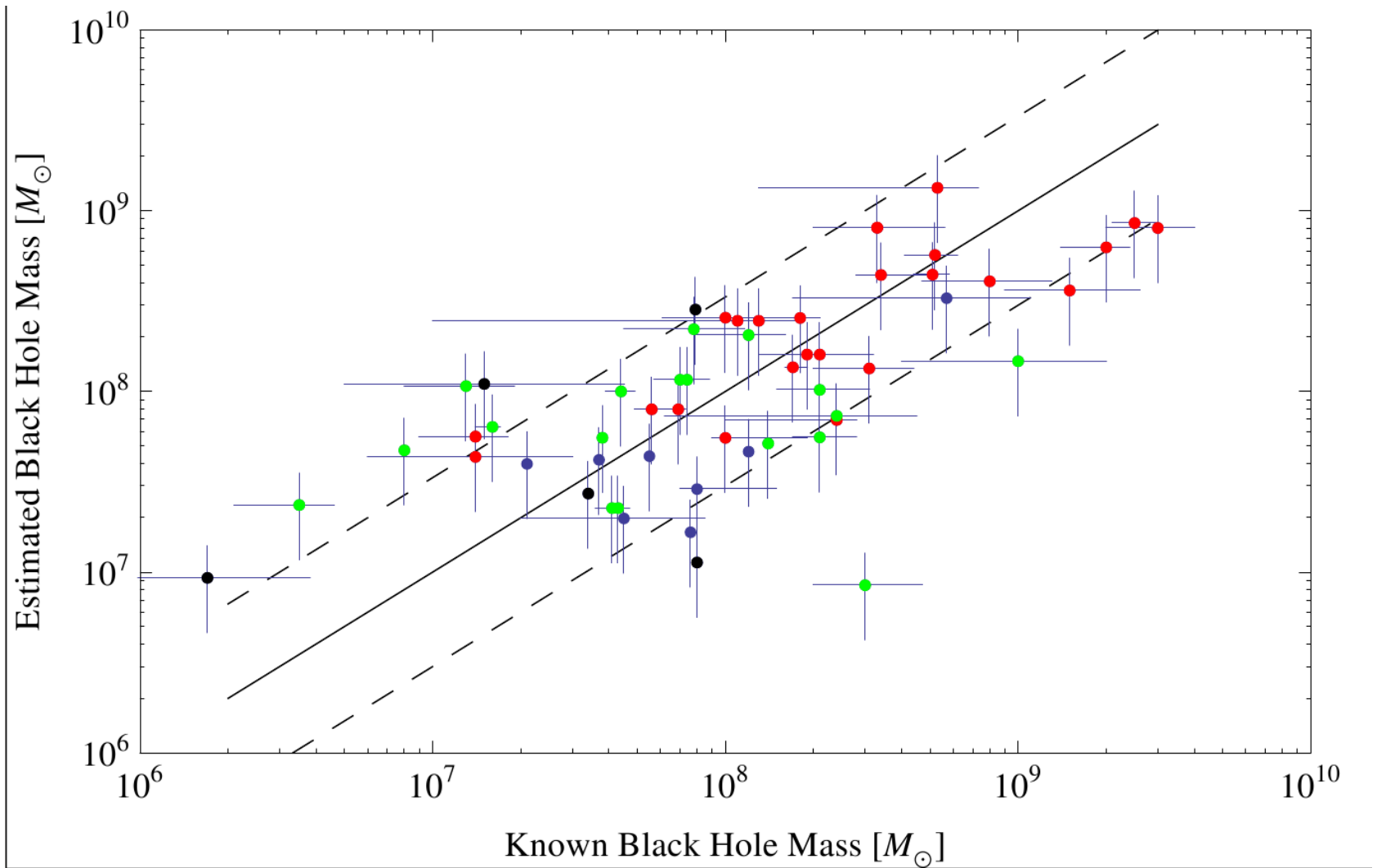
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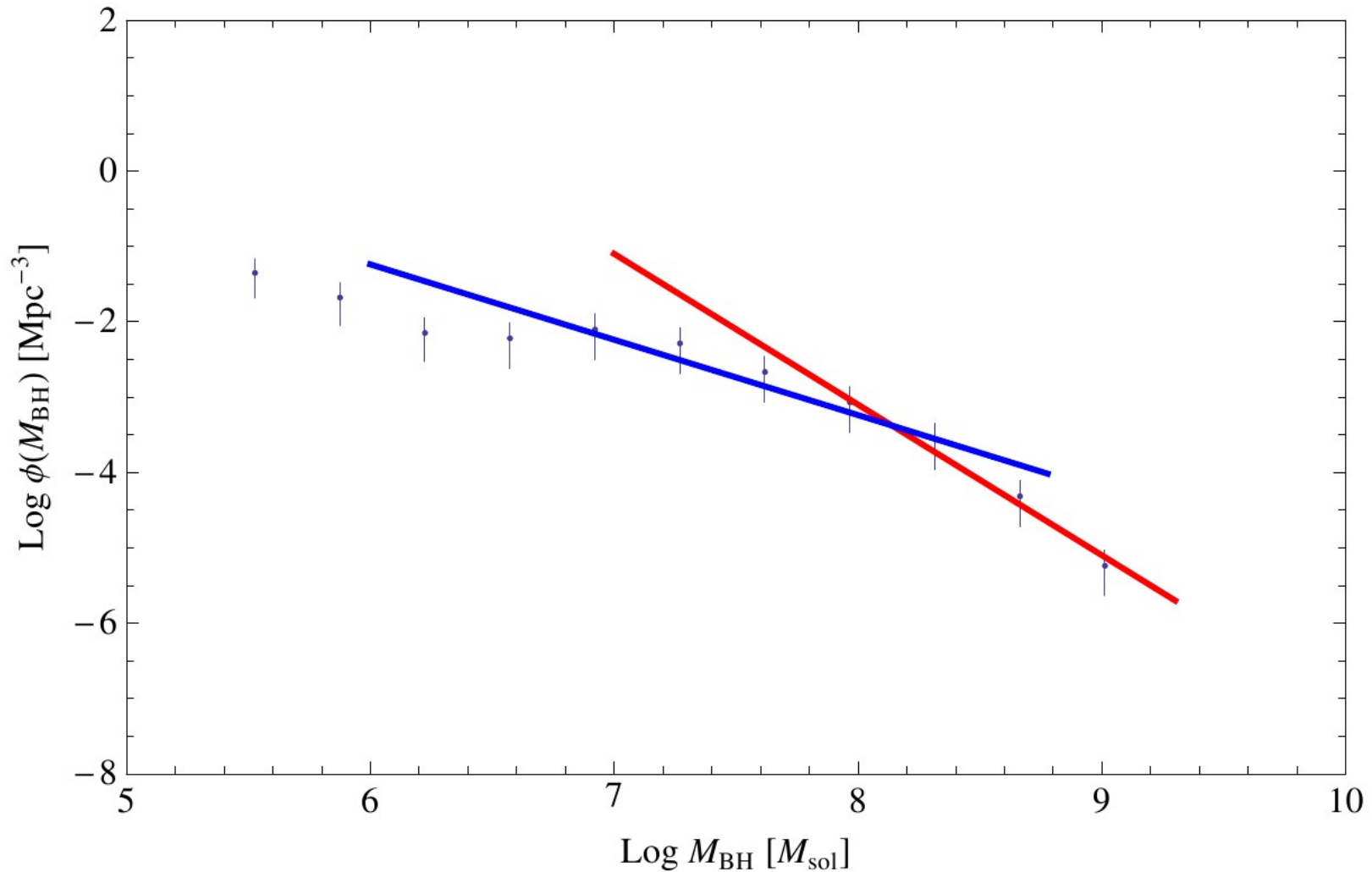
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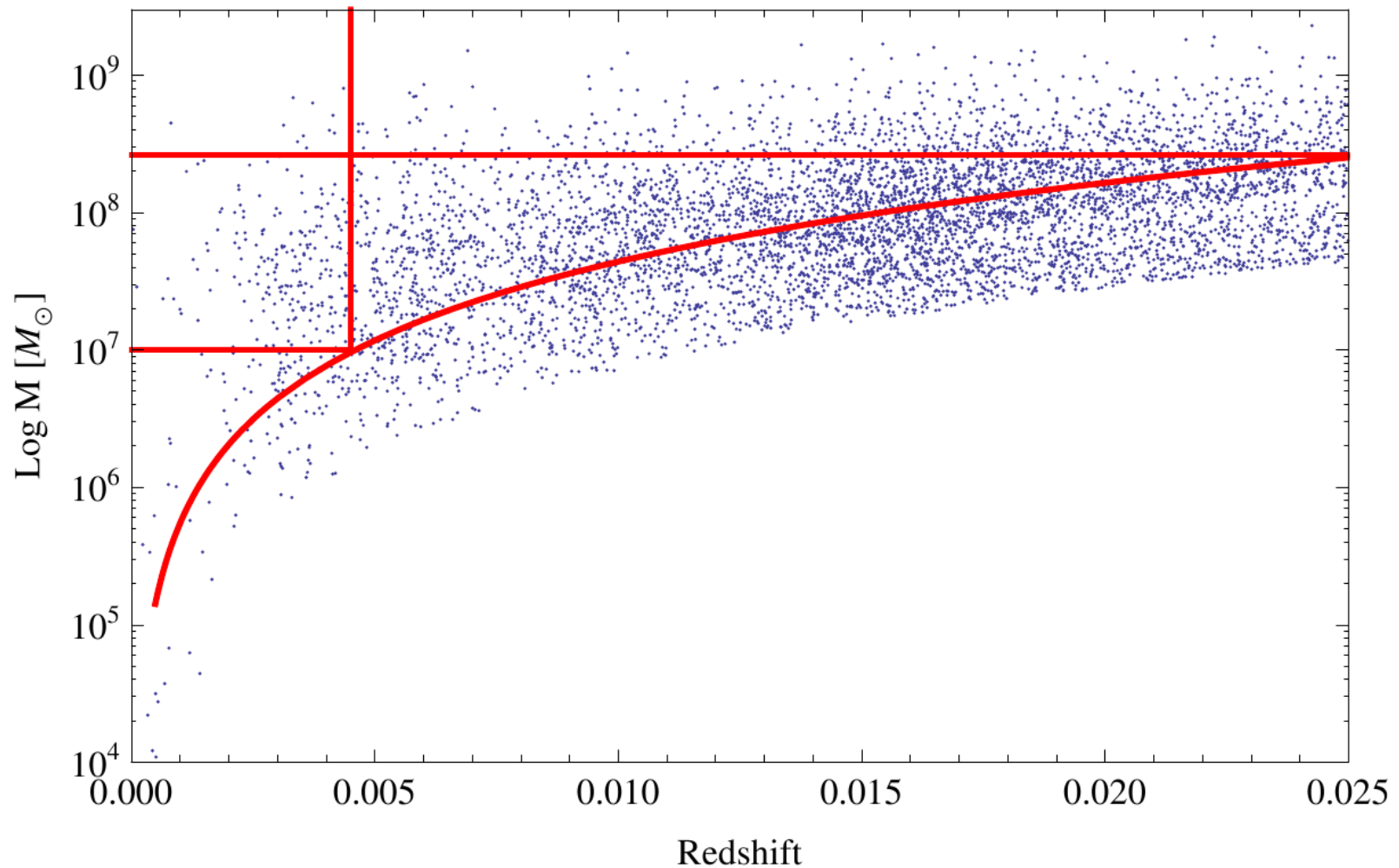


- More info



Integral mass function corrected for Hubble type sampling, 2928 objects, the slope of the lines is: red line -2.0 fitting $> 10^8$ M_{sol} , and blue line -1.0 fitting between 10^7 M_{sol} and 10^8 M_{sol} .

- More info



Plot of Mass over redshift for the massive black hole catalog with the selection curve for elliptical galaxies in blue. This also shows the stepwise selection procedure with the two most extreme cases, the lowest mass, and the highest redshift, also in red.