





SKA Pathfinderers

Hans-Rainer Klöckner



Slides stolen from
Joseph Lazio, Richard Schilizzi, Justin Jonas,
Colin Lonsdale, Leo Blitz, David DeBoer, Daniel Mitchell, ...



path finding



... boldly go where no man has gone before.

QuickTime and a decompressor are needed to see this picture.

But when, why and where we are going ?



When - SKA timeline



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decompressor
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Why - SKA key science drivers



Probing the Dark Ages

- When & how were the first stars formed?

Cosmology and Galaxy Evolution

- Galaxies, Dark Energy and Dark Matter

Strong-field tests of General Relativity

- Was Einstein correct?

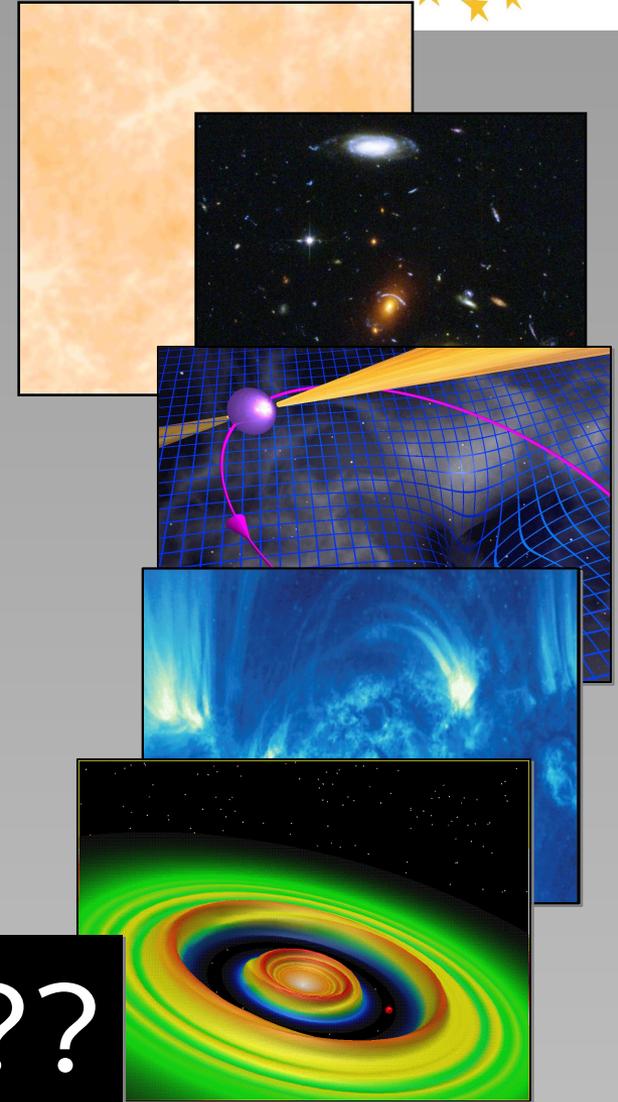
Origin & Evolution of Cosmic Magnetism

- Where does magnetism come from?

Cradle of Life

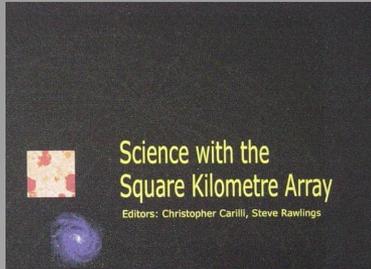
- What and where are the conditions for life?

Exploration of the Unknown



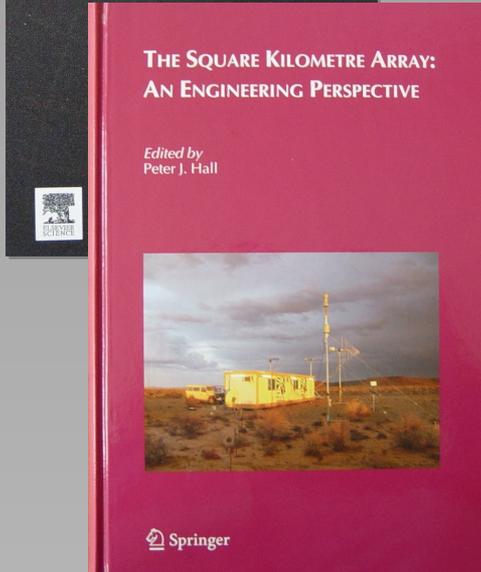


SKA literature



Science with the Square Kilometre Array

eds: C.Carilli, S.Rawlings,
New Astronomy Reviews, Vol.48, Elsevier, Dec. 2004



The Square Kilometre Array An Engineering Perspective

eds: P. Hall
Experimental Astronomy, Vol.17, Springer, Nov. 2004

Cosmology, Galaxy Formation and Astroparticle Physics on the Pathway to the SKA

eds: H-R. Klöckner, S.Rawlings, M. Jarvis, A. Taylor
Astron , March 2008

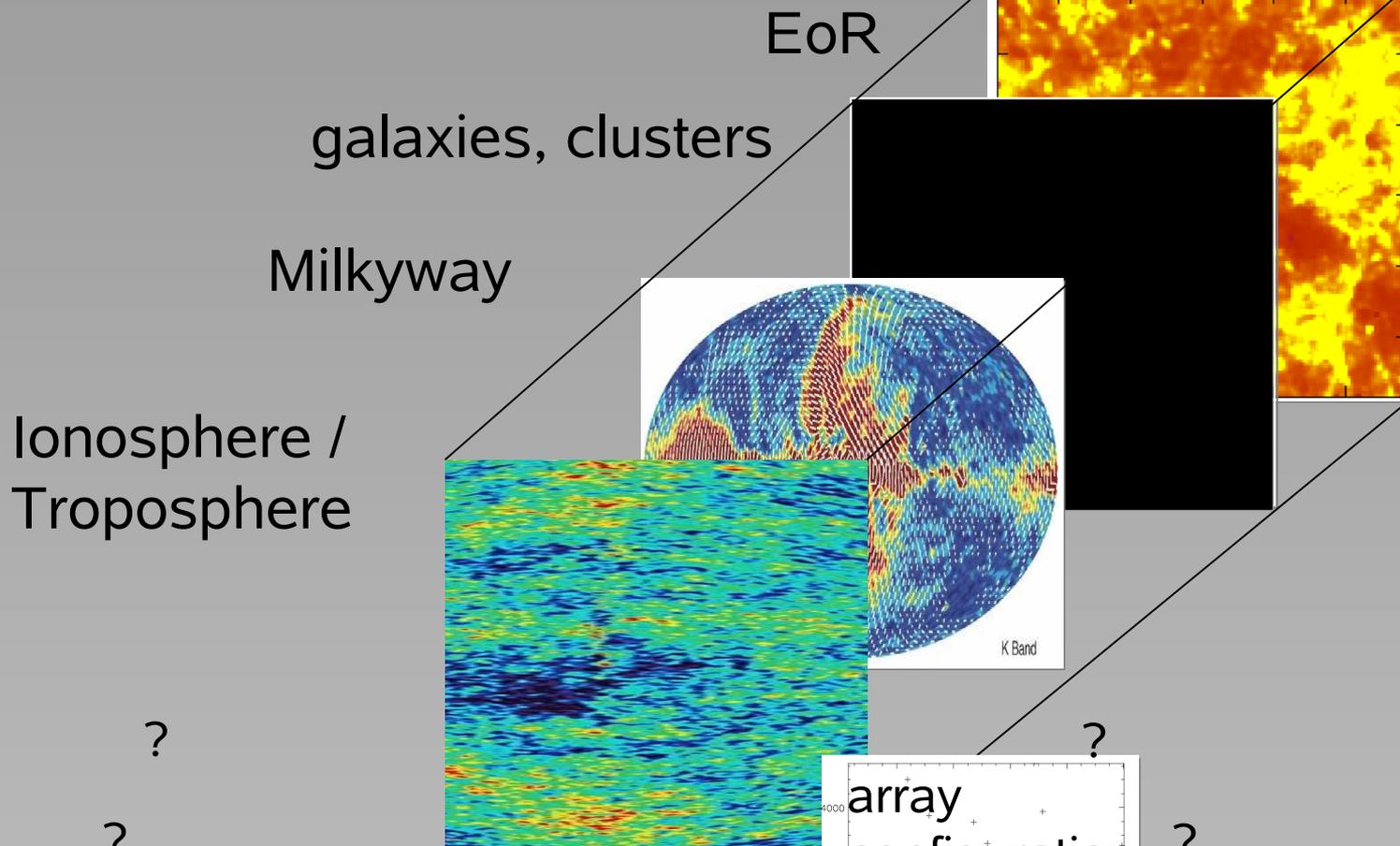
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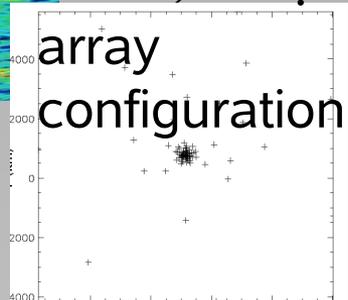
Where - SKA requirements

- sensitivity to detect and image hydrogen in the early universe
→ **very large collecting area**
- **Central concentration of the collecting area** for *optimal detection of hydrogen, pulsars, and magnetic fields*
- fast surveying capability over the whole sky
→ **very large angle field of view and broad frequency coverage**
- capability for detailed imaging of compact objects like active galactic nuclei → **large physical extent**

What do we want to understand



QuickTime and a decompressor are needed to see this [unclear]



Inventory of Telescopes



WSRT



VLA

ATCA



GMRT

		[m ²]	dish [m]	SKA [%]
ATCA	6 x 22 m	760	31	0.0
WSRT	14 x 25 m	6872	94	0.7
VLA	27 x 25 m	13253	130	1
GMRT	30 x 45 m	47712	246	5
Arecibo	1 x 305 m	73061	305	7
LOFAR				< 17



Arecibo

Not to forget the high freq. arrays (e.g. ALMA, SMA)



Shopping list of a “NEW” Interferometer



Do we need to
redo all ?

Telescopes
Software
Hardware

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decompressor
are needed to see this picture.

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SKA station

0.07 – 0.3 GHz (low-band), 0.3 – 10 GHz (mid-band), 10 – >25 GHz (hi-band)

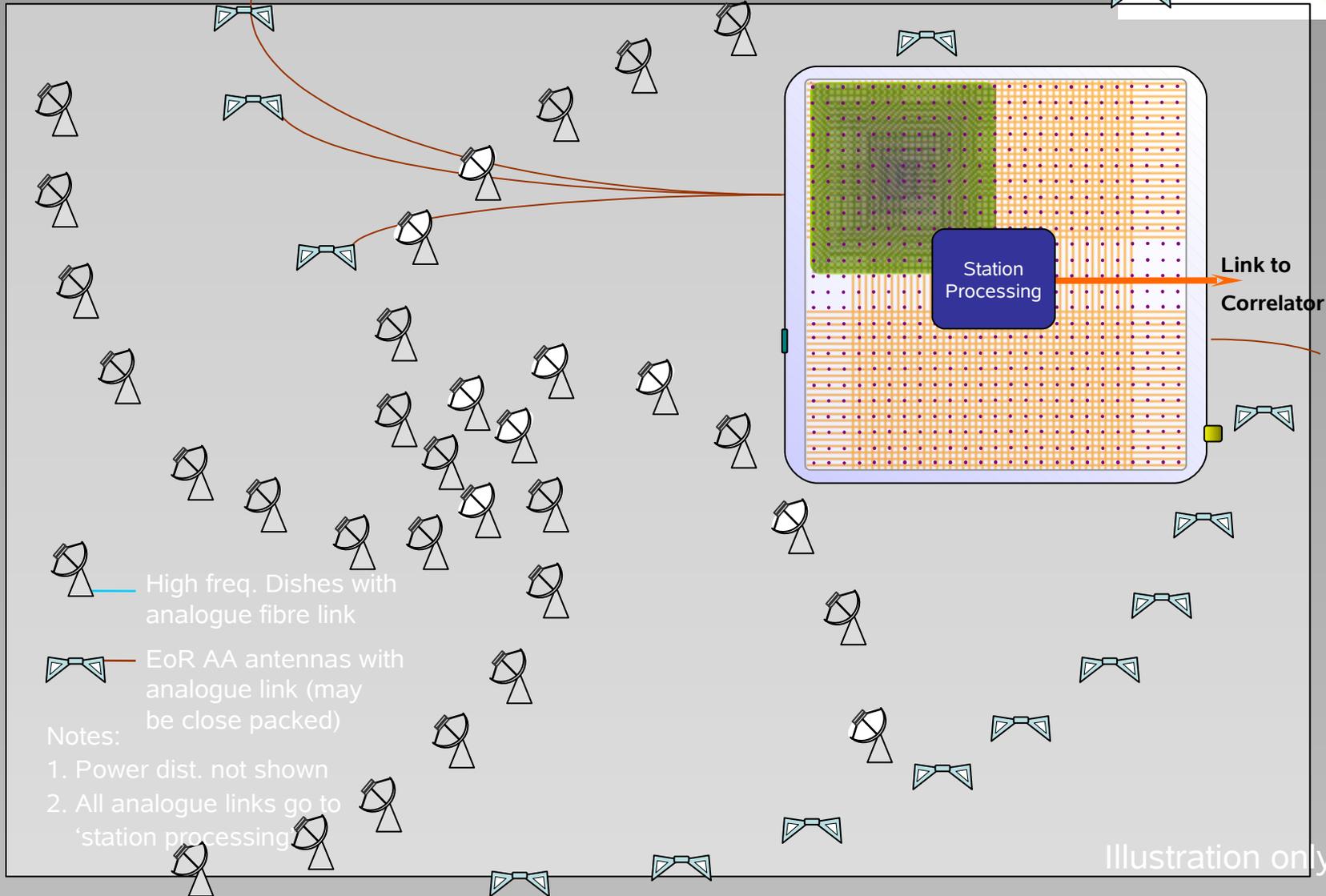


Illustration only



Dishes & Single Pixel Feeds



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Dishes & Phased Array Feeds



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Spare aperture arrays



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Spare aperture arrays



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Shopping list of a “NEW” Interferometer



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current software

Dynamic range

- Antenna sidelobes [bright sources]
- Pointing
- Beam calibration

Ionosphere Calibration

wide bandwidth deconvolution



post

processing

Calibration & Imaging

continues updating
a sky model

new algorithm for
parallelized synthesis



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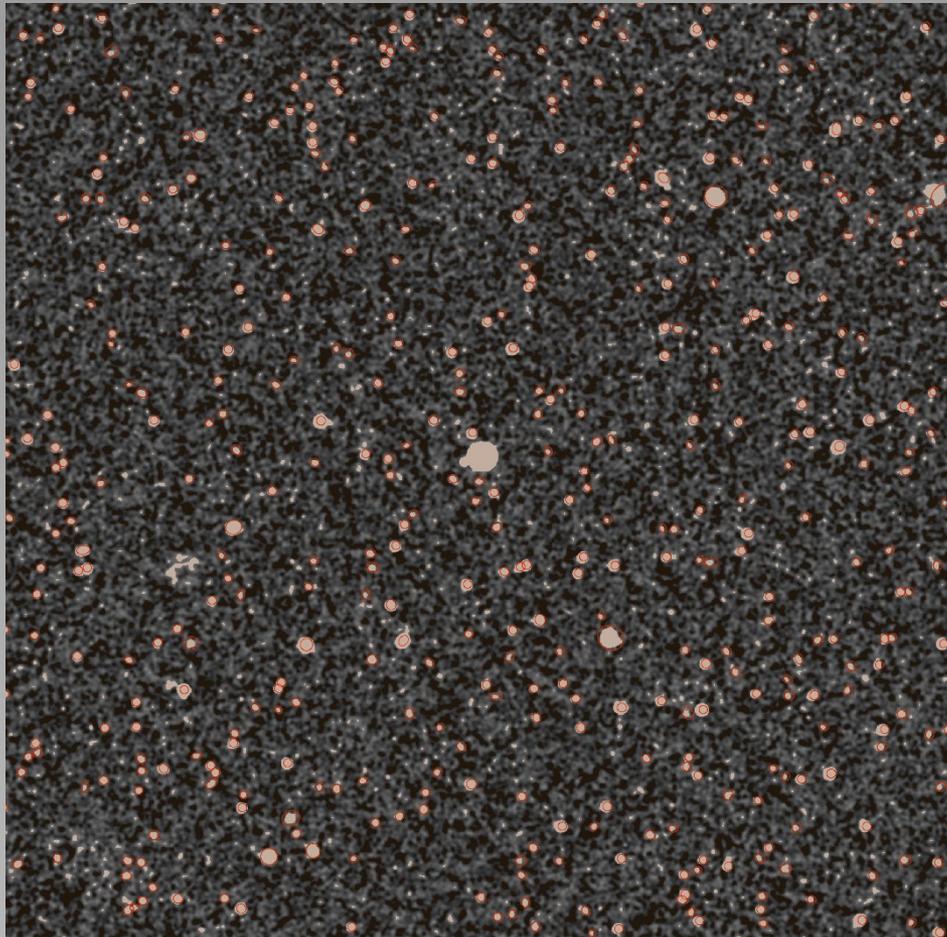
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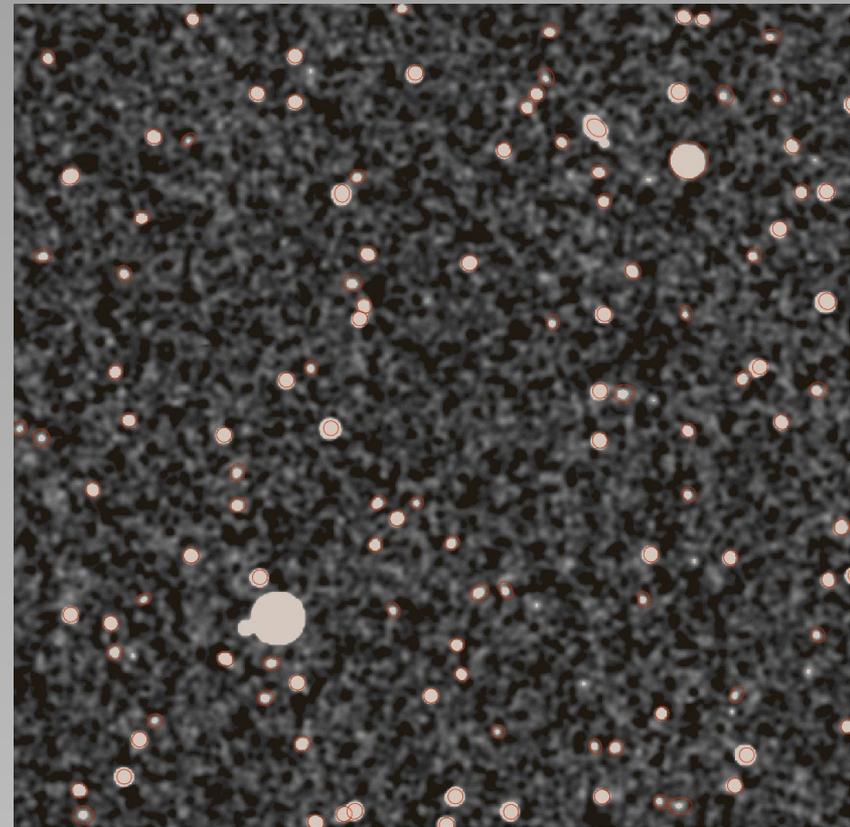
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analysis of the continuum sky



SAD (AIPS), Source extractor,
Miriad



diffuse & extended Structure is the challenge

Source HI Extraction

R. Auld (Cardiff)



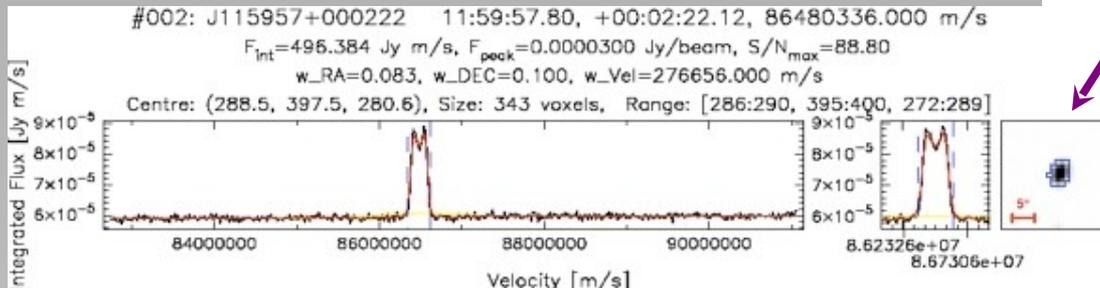
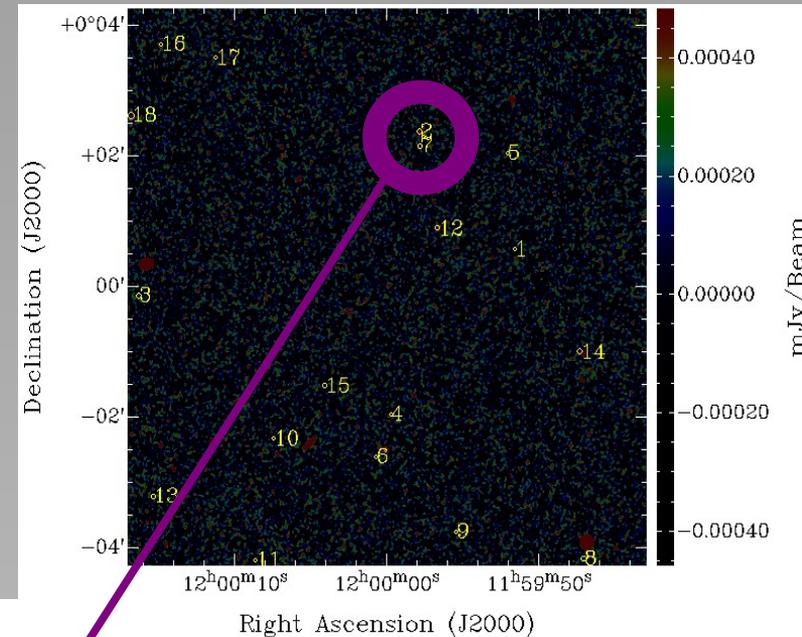
MapMaker

HI cube

Duchamp

Duchamp

- ATNF (M. Whiting)
- Locates sets of contiguous voxels above some threshold
- Uses spectral, spatial or wavelet smoothing for enhancement
- Very quick (30 minutes for 512 x 512 x 1024 cube on a laptop)
- VO-compliant
- Memory and process intensive
- **Completeness issues (misses high S/N sources)**





Sizes of Catalogue



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Hardware

a lot of it



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SKA - Pathfinder



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LOFAR
Ralph Wijers

also touch
LWA
FAST

- location
- hardware in place
- science
- technical development

a NOT complete view



Global location



LOFAR
/

ATA

LWA

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NVSS - VLA

MeerKat

MWA

ASKAP

Galactic Centre

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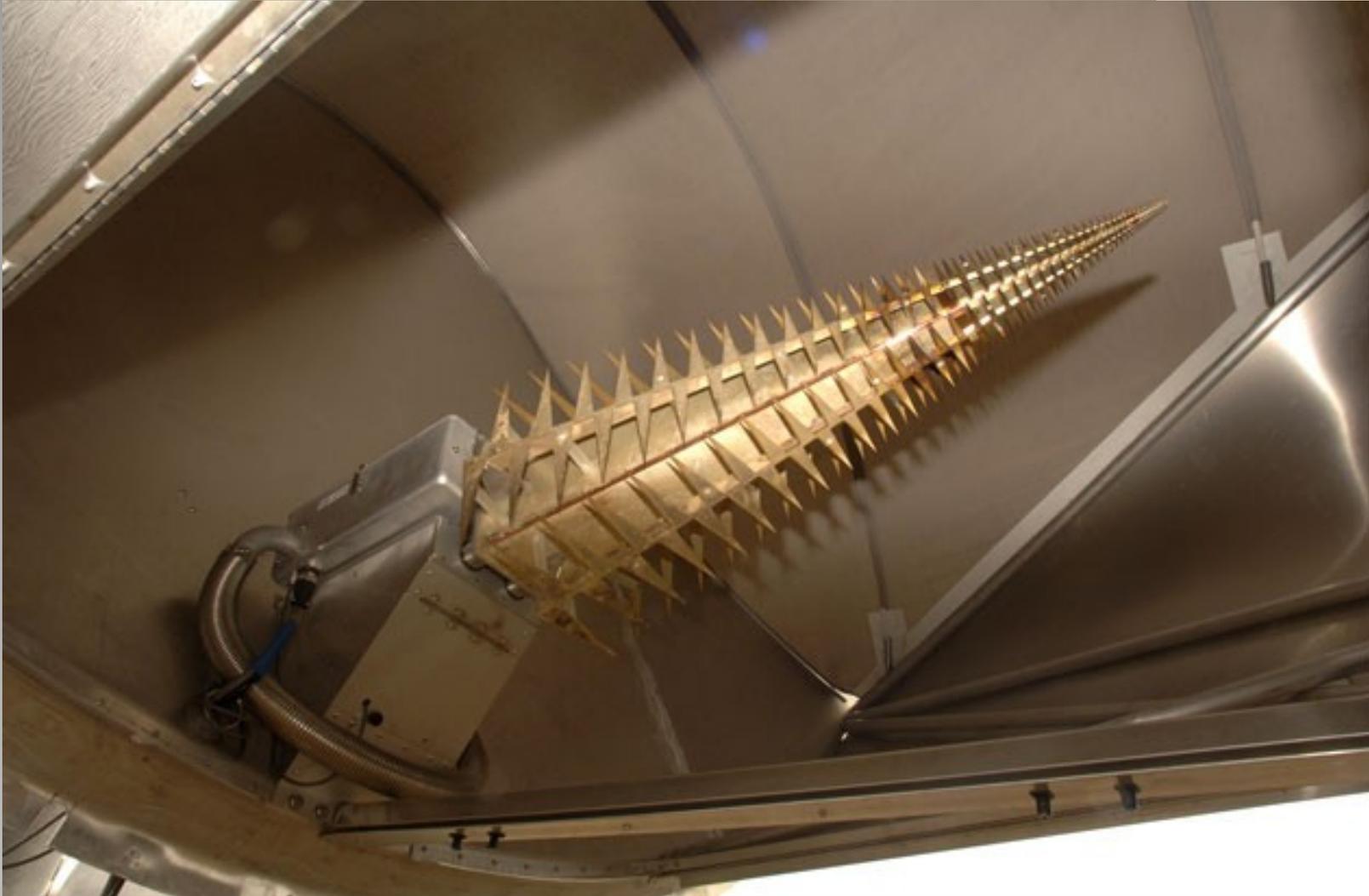


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MWA Science



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path finding



... boldly go where no man has gone before.

Ahhhhh my brain hurts are
we too ambitious ?

NO

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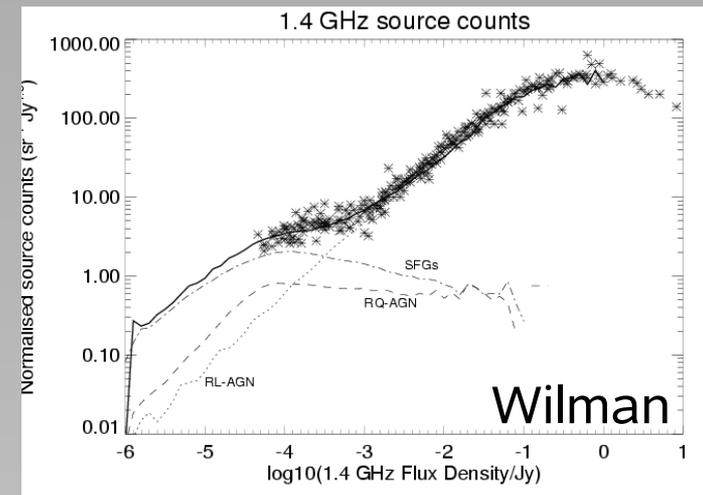
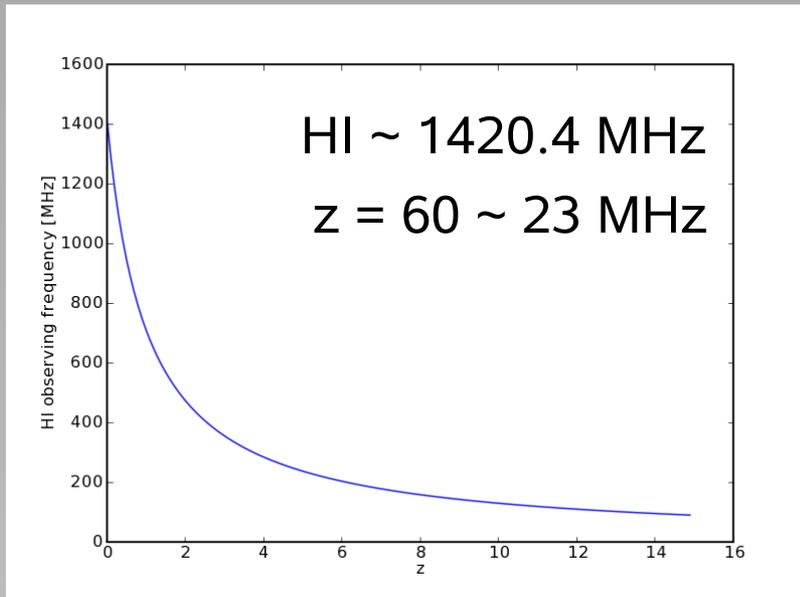
Putting this a little bit together



LWA	20 MHz - 80 MHz
MWA	80 MHz - 300 MHz
LOFAR	30 MHz - 80 MHz & 120 MHz - 240 MHz
ASKAP	700 MHz - 1.8 GHz
KAT	500 MHz - 2.5 GHz
ATA	500 MHz - 11.2 GHz

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Kellerman





Science with Pathfinder



Line observations:

need to wait until they have superior $T_{\text{sys}}/A_{\text{eff}}$ ratio
but opportunity to do Galactic research

Continuum observations:

broader bandwidth $T_{\text{sys}}/A_{\text{eff}}$ ratio

Pulsar - high frequency

Transient

EoR

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Ken Kellerman



Science with Pathfinder



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... and much more !!!

from Justin Jonas



2.0 ASKAP 30
1.4 ASKAP 45
1.2 Meerka 50
0.76 Meerkat 80
5.8 ATA 42
0.69 ATA 350





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