



# MERLIN background e-MERLIN

- Project
- Science



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# MERLIN: background (1)

Jodrell Bank pioneered longbaseline interferometers (>100km) with small remote telescopes and radio links

- Driven by quest to establish nature and size of 'radio stars' : 1950's & 60's
- Key technical development was 'phase stable link' (1973)
  - Most of information through phase
  - Allows coherent integration to make deep images





# MERLIN: background (2)

#### By 1980 developed into MERLIN

- Network of 6 telescopes; 127km
- Extended to 217km with telescope at Cambridge (1990)
- Motivated by high resolution (50 150 mas at cm wavelengths); real time
- 7x resolution of VLA: can resolve
  - Novae; SNR shells in M82
  - Starforming galaxies at z~1
- Key technical advances
  - Affordable data links with 30 MHz b/w
  - Clock distribution (few ps)
- Sensitivity limited by link bandwidth
  - ~ 50 uJy/beam at 5 GHz







#### **Optical fibres**

- Optical fibres allow ~Tb/s per fibre over 100km without amplification already used in ATCA, GMRT key part of ALMA, EVLA
- Available as part of national and international commercial and research communications infrastructures
- Can support GHz bandwidths for e-MERLIN and Gb/s realtime operation for EVN





# e-MERLIN Project

- Funded from 2002: Univ Manchester, NWDA, PPARC/STFC, Cambridge, UMIST, LJMU. £8M
- Tightly focused on sensitivity upgrade using fibre links for 2 GHz bandwidth
- Design goals
  - 2 GHz bandwidth (in each pol)
  - New C-band: 4-8 GHz Rx suite
  - Upgrade L-band: 1.3-1.7 GHz
  - Upgrade K-band: 22-24 GHz
  - Rapid (1min) change between bands
  - No compromise in phase stability
  - Correlator: Wide-field (<0.5MHz) &
    - spectral line (<0.5 kHz)
- Low-cost
  - share development (ALMA/EVLA/e-MERLIN DTS)
  - Re-use some components



### e-MERLIN: Progress

- Receivers: done
- IF & electronics: prototype
- Fibre network:
- Correlator: DRA prototype
- Software: in pro
- First fringes with correlator: Mid (



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## WIDAR Correlator

- Developed by DRAO, Penticton (Canada) and NRAO [for EVLA]
  - 2 x 2 GHz wide band sampled at telescope (2x 4 Gs/s with 3 bit samples) and transmitted using 3 x 10Gb/s optical wavelengths with custom transmission equipment
  - Correlator splits input to 16 x 128 MHz sub-bands using digital filters
  - Sub-bands correlated using 16 x 64 x 2048 complex-lag correlator chips
    - 0.25 MHz channel resolution for 4 polarisations at maximum bandwidth
    - KHz resolution for spectroscopy, Hz resolution for radar



### Capabilities

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  - 150, 40, 10 mas resolution at L[20cm],C[6cm],K[1,3cm]
  - ~2 uJy sensitivity in typical runs
    - <uJy in deep fields</li>
    - ~30 uJy in ~1 min
  - Wide fields
    - Out to HPBW of 25-m [9,30 arcmin]
  - Spectroscopy
    - 16 placeable sub-bands; >512 channels/pol; recirculation
    - Can mix/trade bandwidths; no. of channels, polarisations
  - Much improved aperture coverage
    - Via frequency coverage
    - May help snapshots too
  - Spectral mapping
    - 1.3-1.7; 5-7/4-8 GHz
  - Polarization (L,R  $\rightarrow$  IQUV)
  - Astrometry
    - Goal is < 1 mas wrt ICRF

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#### e-MERLIN: Science





#### e-MERLIN: Science

- MANCHESTER
  - Key topics:
    - Formation processes of stars
      - Jets, winds, disks
    - Activities in nearby galaxies
      - Black holes, supernovae, starclusters
    - Distant galaxies
      - Starformation & AGN
    - Gravitational lenses
      - Dark matter distribution
  - All require >100 km baselines







## Using MERLIN

- Open facility
- Propose using Northstar tool
- Deadlines Mar & Sep
- Limited observations in 2008
- www.merlin.ac.uk



# Using e-MERLIN

- MANCHESTEI 1821
- Powerful new combination of high resolution and sensitivity
- Data analysis similar to MERLIN, EVLA
- Open facility
- Legacy programme for large projects www.merlin.ac.uk/e-merlin\_legacy.html Projects being designed now, teams open
- Commissioning & early science 2008
- Full operations 2009