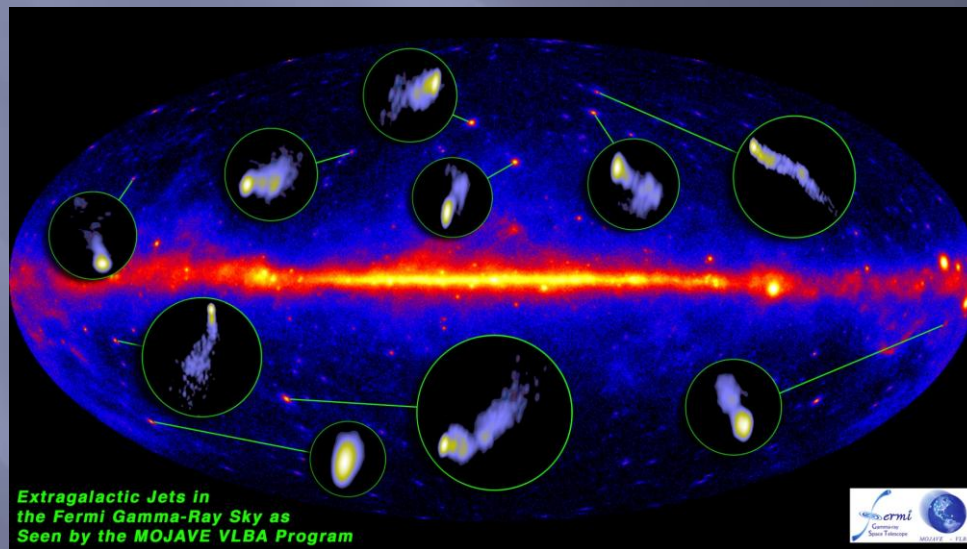
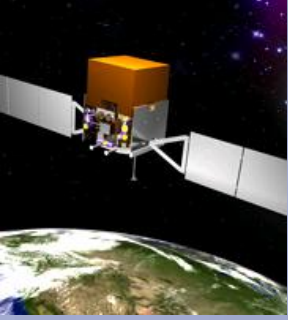


# P5: BEAMING PROPERTIES OF GAMMA- RAY BRIGHT BLAZARS



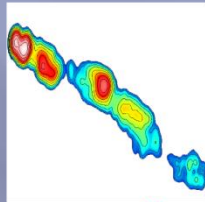
# Data



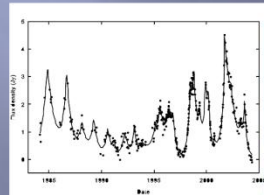
VLBA



Apparent jet speeds



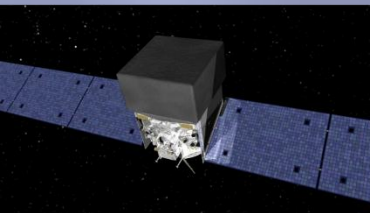
Metsähovi



Doppler factors from mm-variability



Fermi



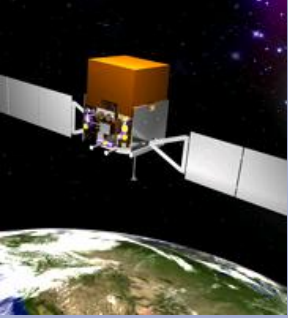
Gamma-ray detections

Intrinsic jet properties:

$$\Gamma, \theta, \theta_{\text{src}}$$

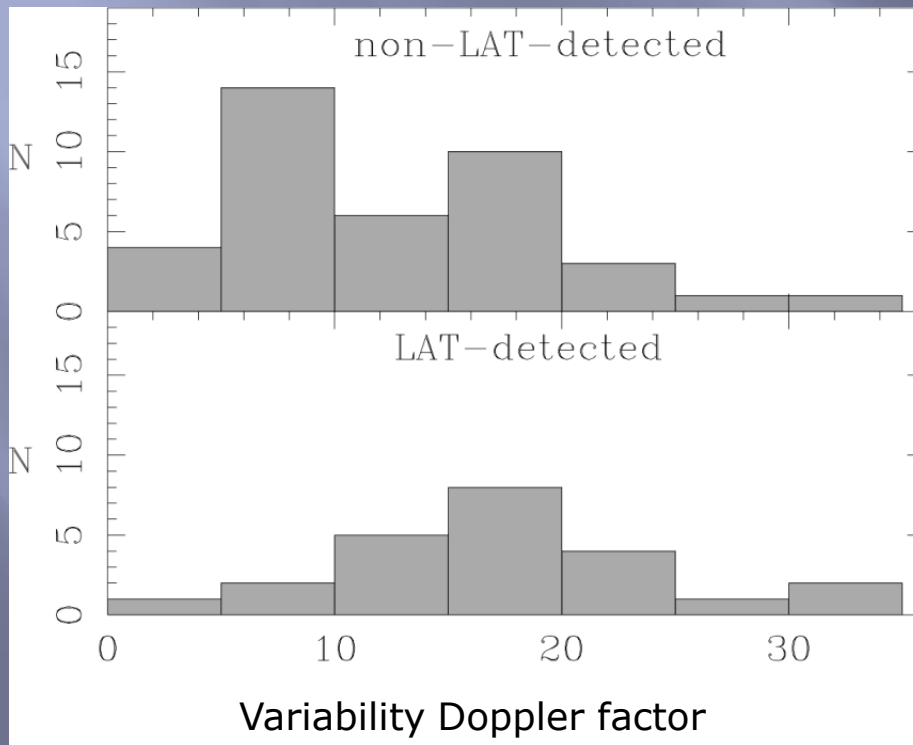
- Apparent jet speeds from MOJAVE survey (*Lister et al. 2009*)
- Doppler factors from the long term mm-monitoring at Metsähovi (*Hovatta et al. 2009*)
- Allows calculation of bulk Lorentz factors and viewing angles in the observer's frame and in the comoving frame of the source
- Sample: 62 blazars, a sub-set of the complete flux-density limited MOJAVE sample

(Radio) jet properties compared with Fermi LAT 3-month list of bright gamma-ray sources.  
(*Savolainen et al. 2010, A&A, 512, 24*)



# Results

Bright  $\gamma$ -ray blazars detected by *Fermi* LAT during its first 3 months of observations have on average higher variability Doppler factors ( $P > 99.3\%$ )

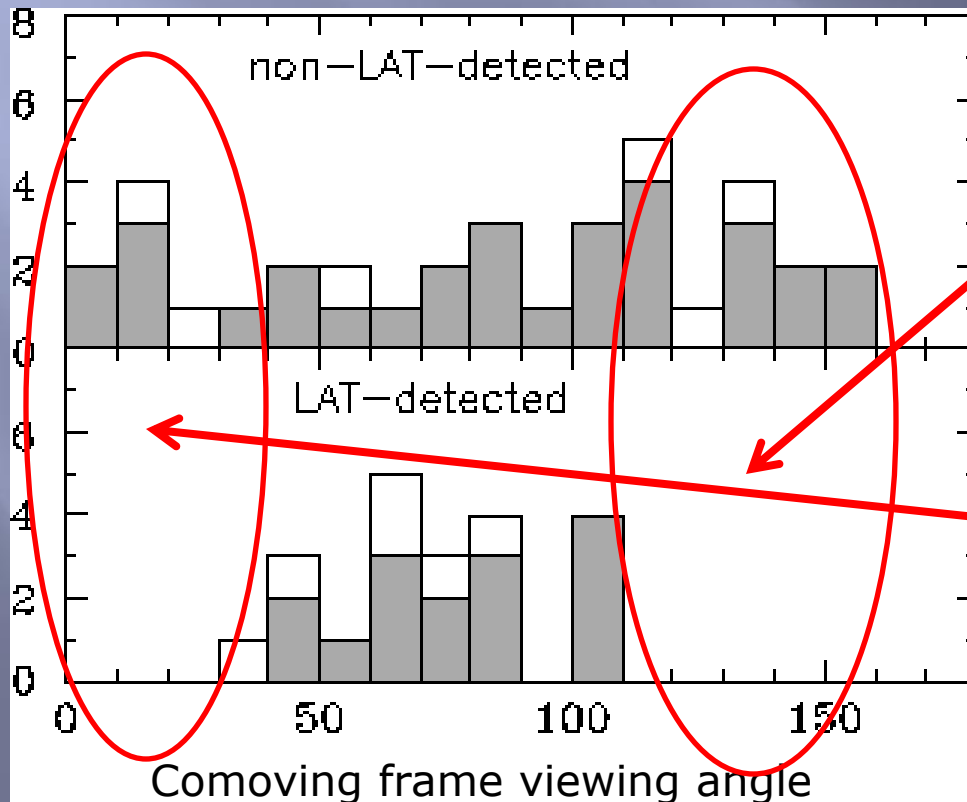


- Agrees with bright gamma-ray blazars having faster apparent jet speeds, wider apparent jet opening angles, and higher core  $T_b \rightarrow$  **more highly beamed**



# Results

... and a narrower distribution of comoving frame viewing angles ( $P > 98\%$ )



- Difference at large comoving frame viewing angles can be explained by beaming

- Difference at small comoving frame viewing angles, if confirmed, requires another explanation. Intrinsic gamma-ray emission anisotropy?