



Max-Planck-Institut
für Radioastronomie



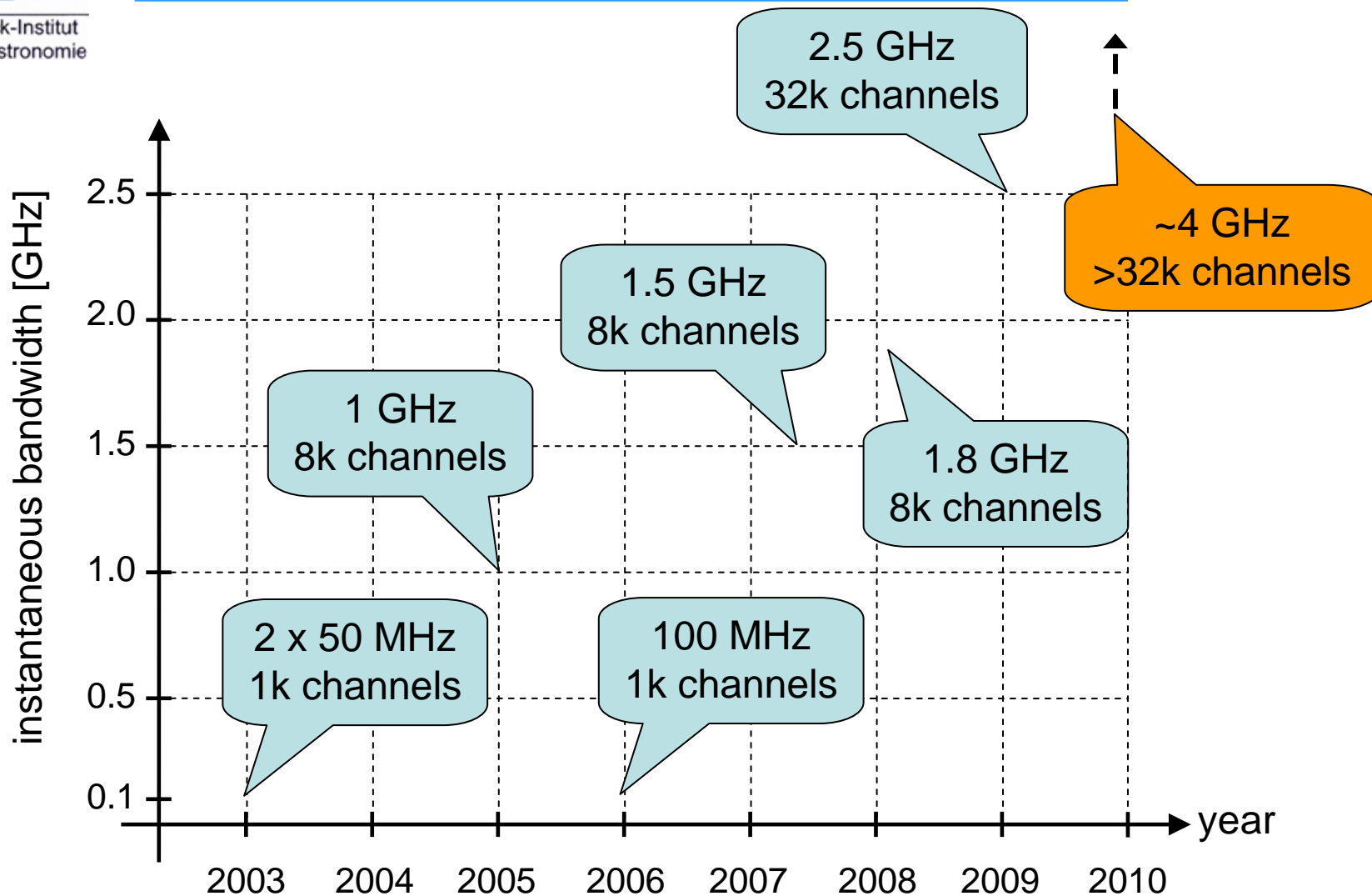
Fast Fourier Transform Spectrometer (FFTS)

Bernd Klein

*Max-Planck-Institut für Radioastronomie, Bonn
- Germany -*



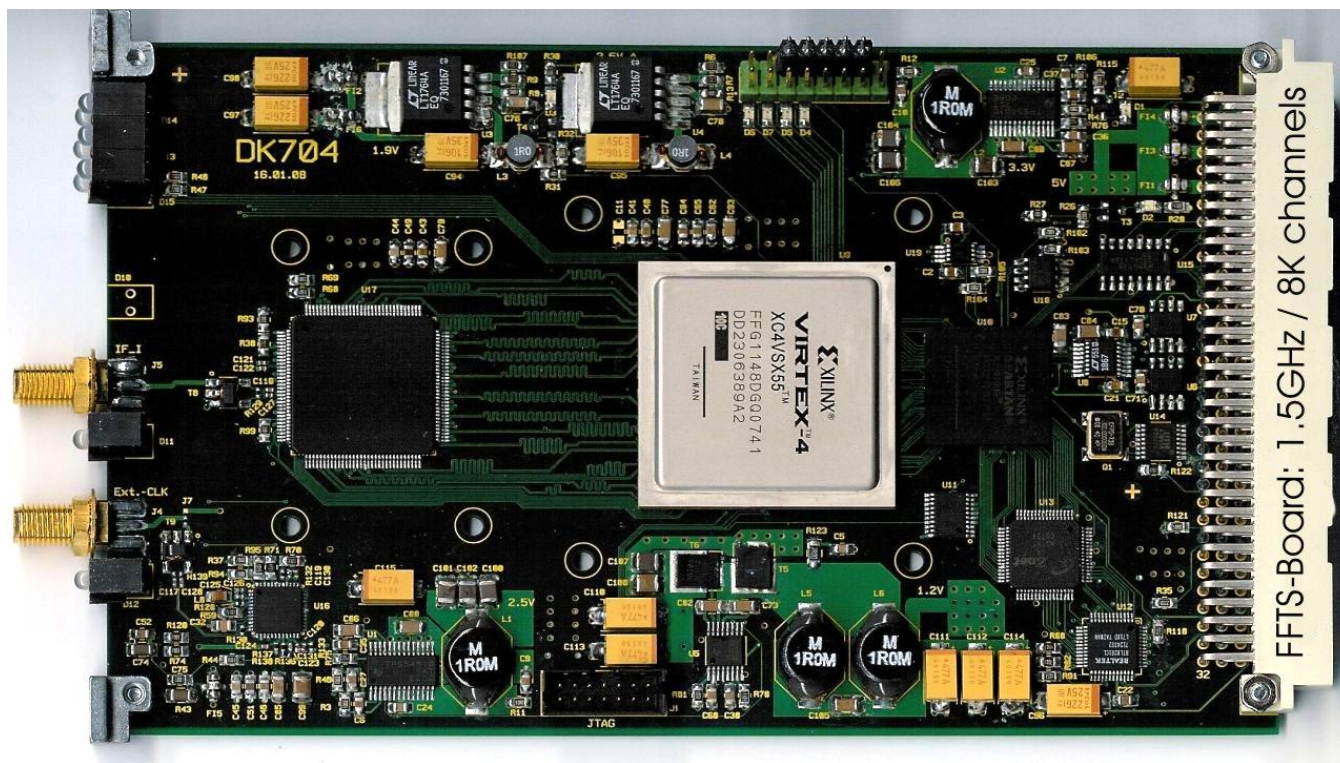
**Improvement in bandwidth by a factor
of 50 over 6 years or >400 MHz per year!**





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FFTS :: The MPIfR-Board



IF input
(0 - 3 GHz)

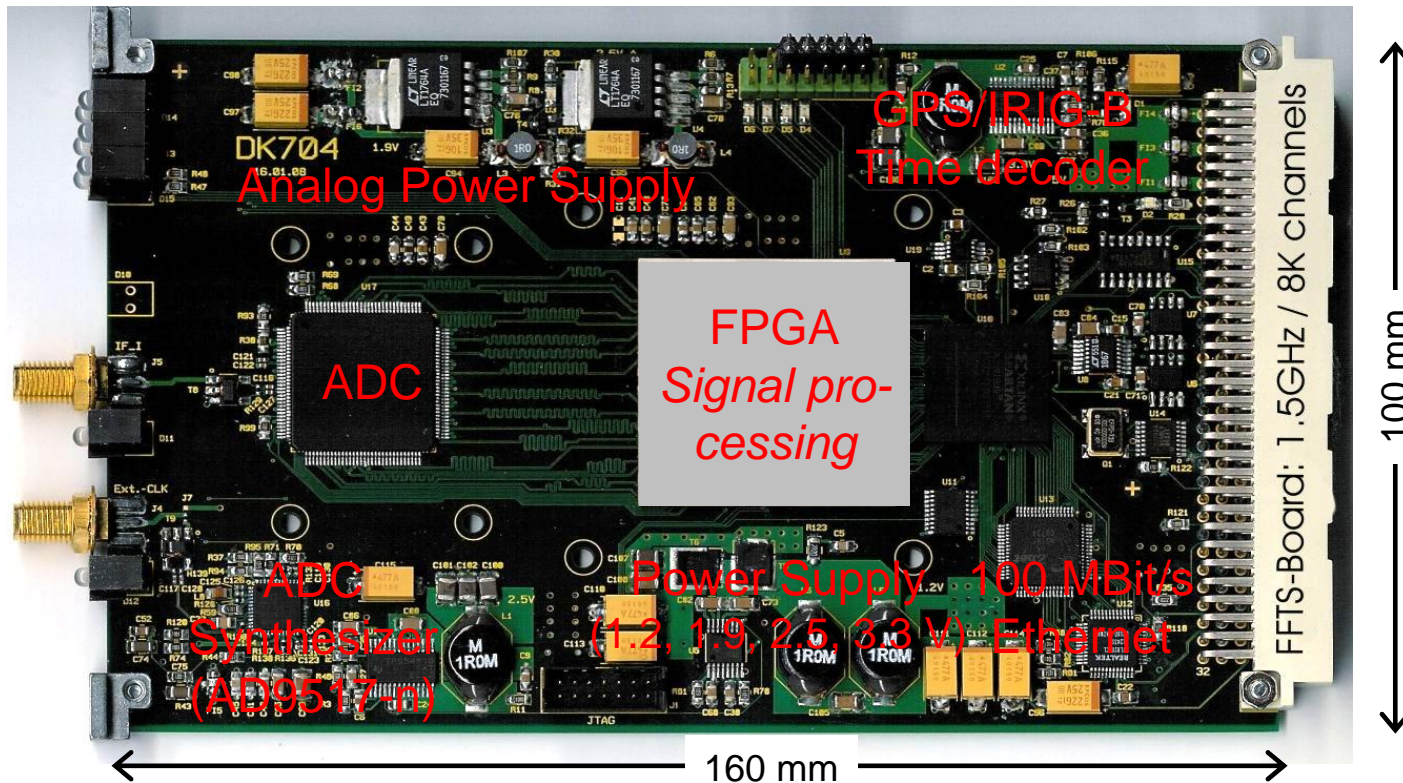
5 Volt
←

Data
→

- Instantaneous bandwidth: 0.1 – 1.8 GHz
- Spectral resolution @ 1.5 GHz: 212 kHz
- Stability (spec. Allan Variance): > 1000 sec.
- Calibration- and aging free digital processing



FFTS :: The MPIfR-Board



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APEX :: Atacama Pathfinder EXperiment

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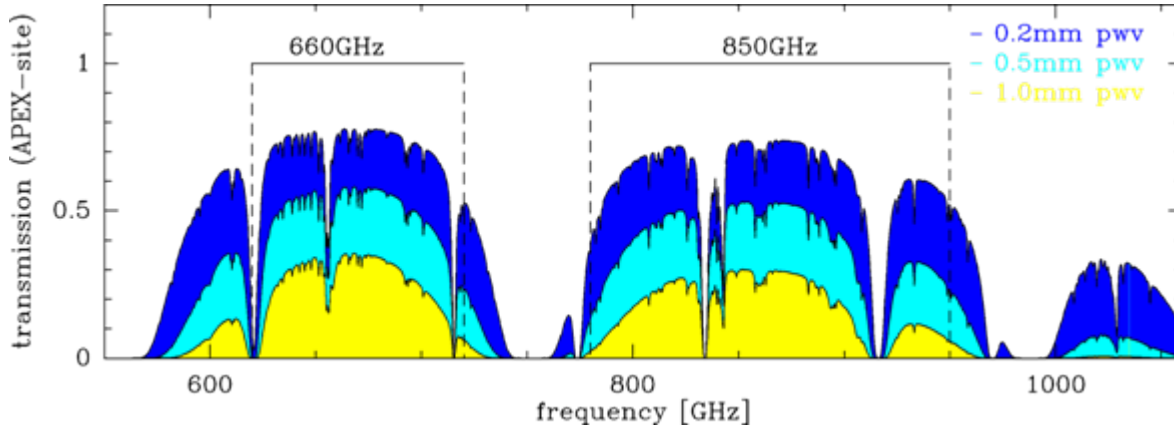
Telescope:

| | |
|---------------------------|--|
| Location: | Llano de Chajnantor, 50 km east San Pedro de Atacama, northern Chile |
| Coordinates: | Latitude : 23°00'20.8" South Longitude : 67°45'33.0" West Elevation : 5105 m |
| Diameter: | 12 m |
| f/D: | 8 |
| Beam width: | (FWHM) 7.8" * (800 / f [GHz]) |
| Main reflector: | 264 aluminum panels, average panel surface rms 5 micron |
| Surface accuracy: | 17 micron rms |
| Pointing accuracy: | 2" rms over sky |
| Mounting: | Alt-Az |
| Receiver cabins: | 2 Nasmyth + 1 Cassegrain |
| Mass: | 125 000 kg |
| Manufacturer: | Vertex Antennentechnik |



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CHAMP+ :: Array-RX 660/850 GHz window



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AFFTS :: Array-FFTS for APEX

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Bandwidth: $32 \times 1.5 \text{ GHz} = 48 \text{ GHz}$ (option 58 GHz)
Spec. channels: $32 \times 8\text{k} = 256\text{k}$ channels @ 212 kHz

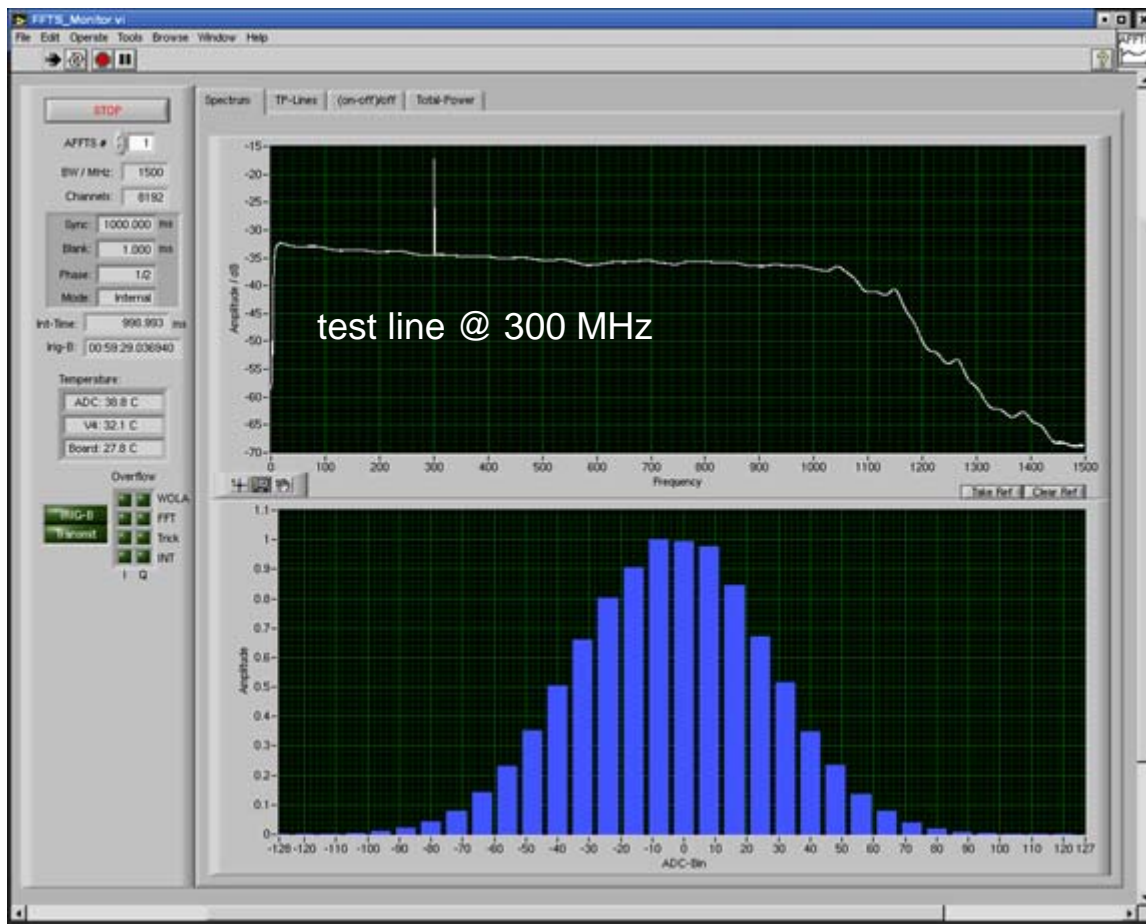


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AFFTS :: Software / Monitoring



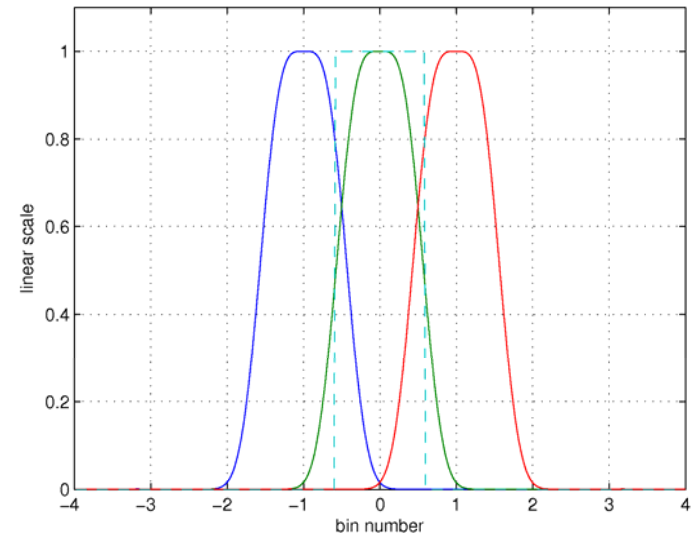
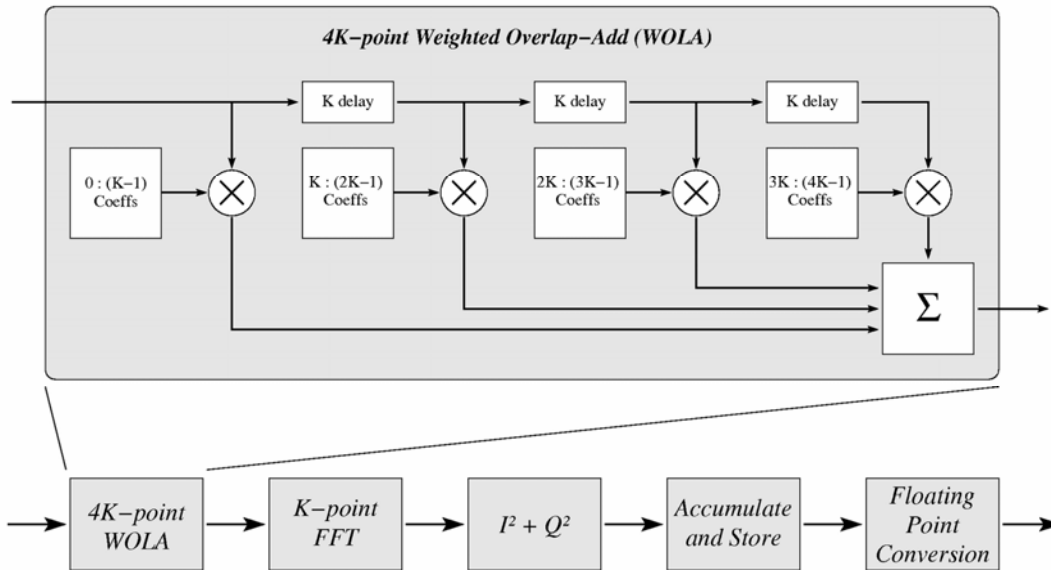
FFTS Control Software:

- multi-threaded software / Linux
- one thread per FFTS-board (smoothing, leveling, RFI-filter, ..)
- communication telescope / FFTS: SCPI protocol (UDP/ASCII), e.g. APEX:AFFTS:band1:cmdBandwidth 1500
- data flow: TCP (header + spectra)
- FFTS Monitor Tool (LabView)



FFTS :: Signal Processing

Unlike the conventional windowed-FFT processing, a more efficient polyphase pre-processing algorithm has been developed with significantly reduced frequency scallop, less noise bandwidth expansion, and faster sidelobe fall-off.



Frequency response of the optimized FFT signal processing pipeline

Equivalent noise bandwidth = 1.16 x frequency spacing



FFTS :: *FPGA configurations*

Today, implemented FFTS board / FPGA configurations are:

- 1 x 1.5 GHz bandwidth, 1 x 8192 spectral channels, ENBW: 212 kHz (default core)
- 1 x 1.8 GHz bandwidth, 1 x 8192 spectral channels, ENBW: 255 kHz

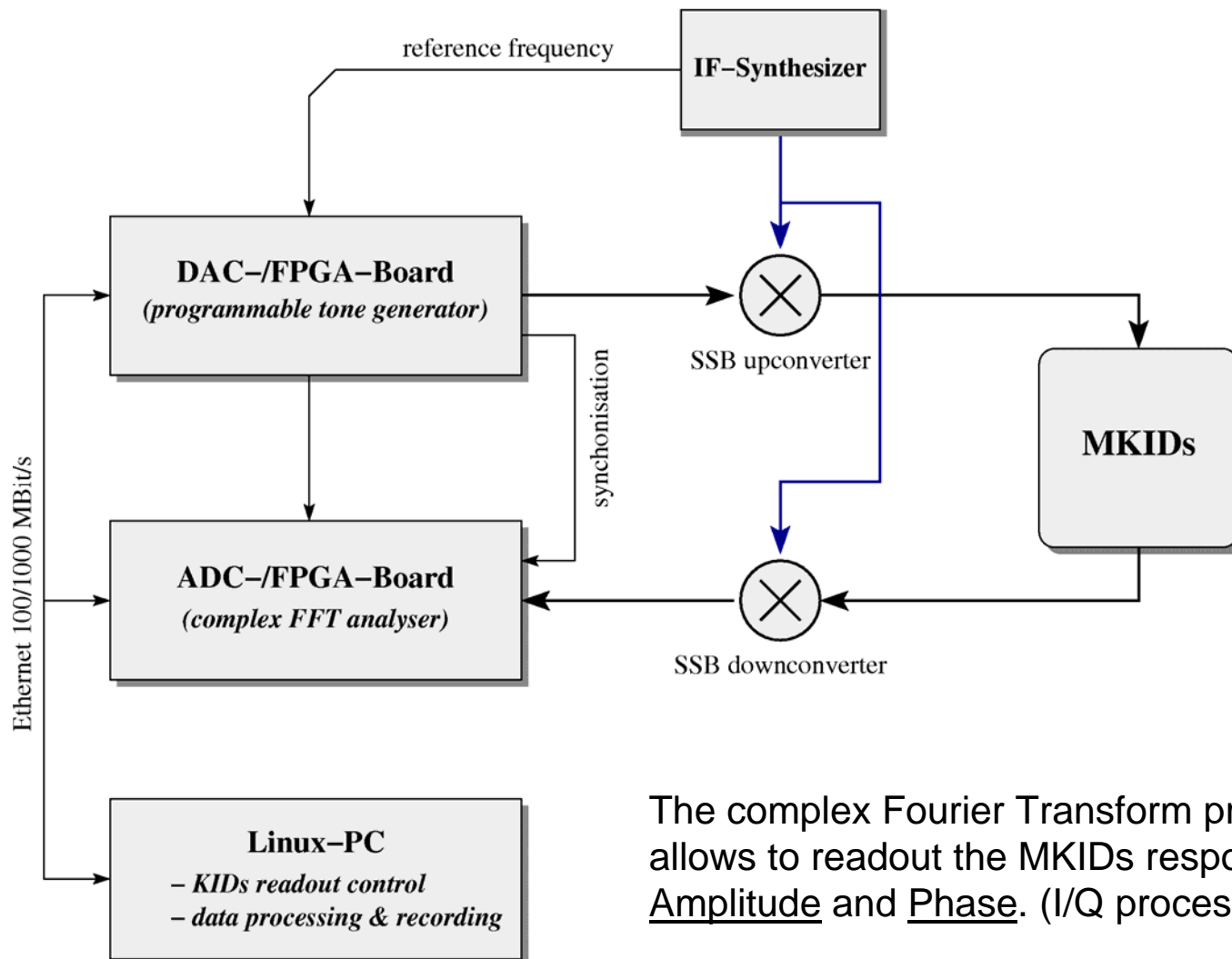
- 1 x 750 MHz bandwidth, 1 x 16382 spectral channels, ENBW: 53 kHz
- 1 x 500 MHz bandwidth, 1 x 16384 spectral channels, ENBW: 35 kHz
- 1 x 100 MHz bandwidth, 1 x 16384 spectral channels, ENBW: 7 kHz
- 1 x 50 MHz bandwidth, 1 x 16384 spectral channels, ENBW: ~4 kHz

- 2 x 750 MHz bandwidth, 2 x 8192 spectral channels, ENBW: 107 kHz
- 2 x 500 MHz bandwidth, 2 x 8192 spectral channels, ENBW: 71 kHz
- 2 x 100 MHz bandwidth, 2 x 8192 spectral channels, ENBW: 14 kHz

The Equivalent Noise Bandwidth (ENBW) is the width of a fictitious rectangular filter such that the power in that rectangular band is equal to the (integrated) response of the actual filter.



MKIDs Readout: *The Principle*



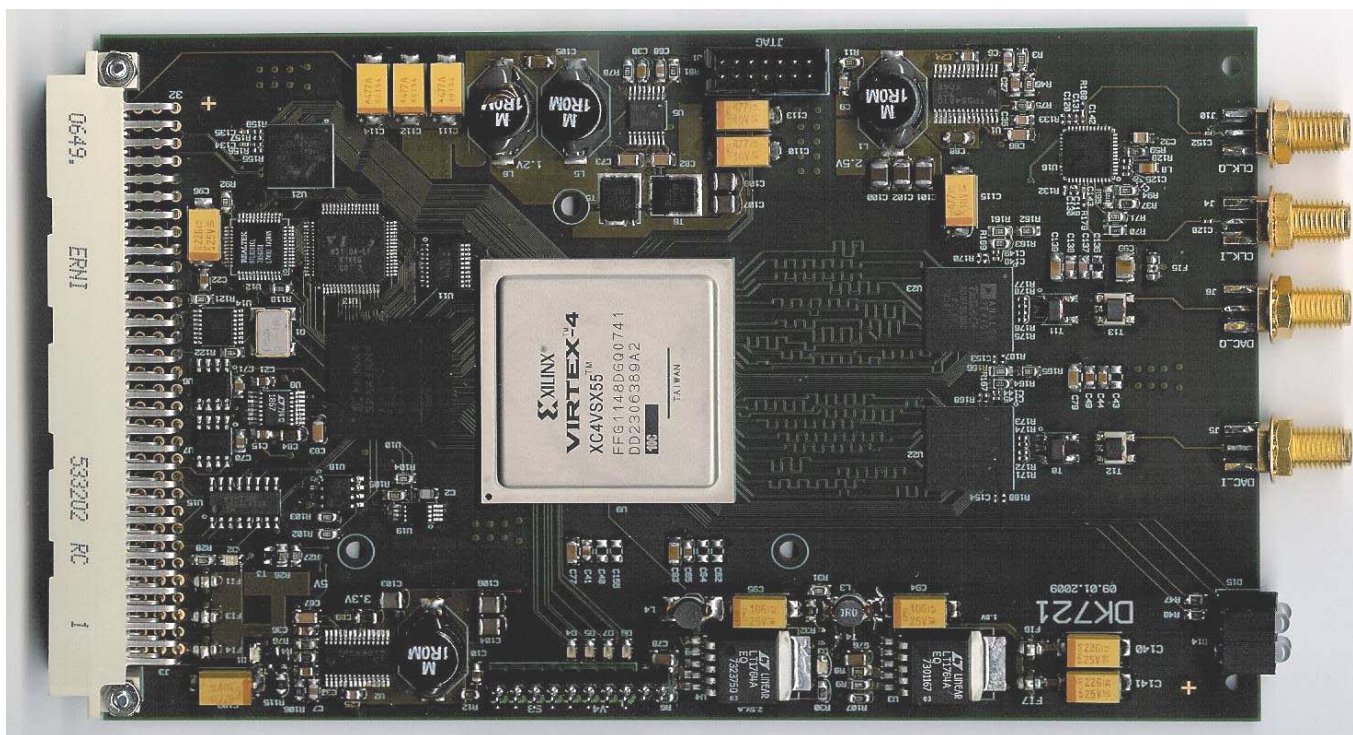
The complex Fourier Transform processing allows to readout the MKIDs response in Amplitude and Phase. (I/Q processing)



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MKIDs Readout: *Prototype Hardware*

DAC-/FPGA-Board – The tone generator



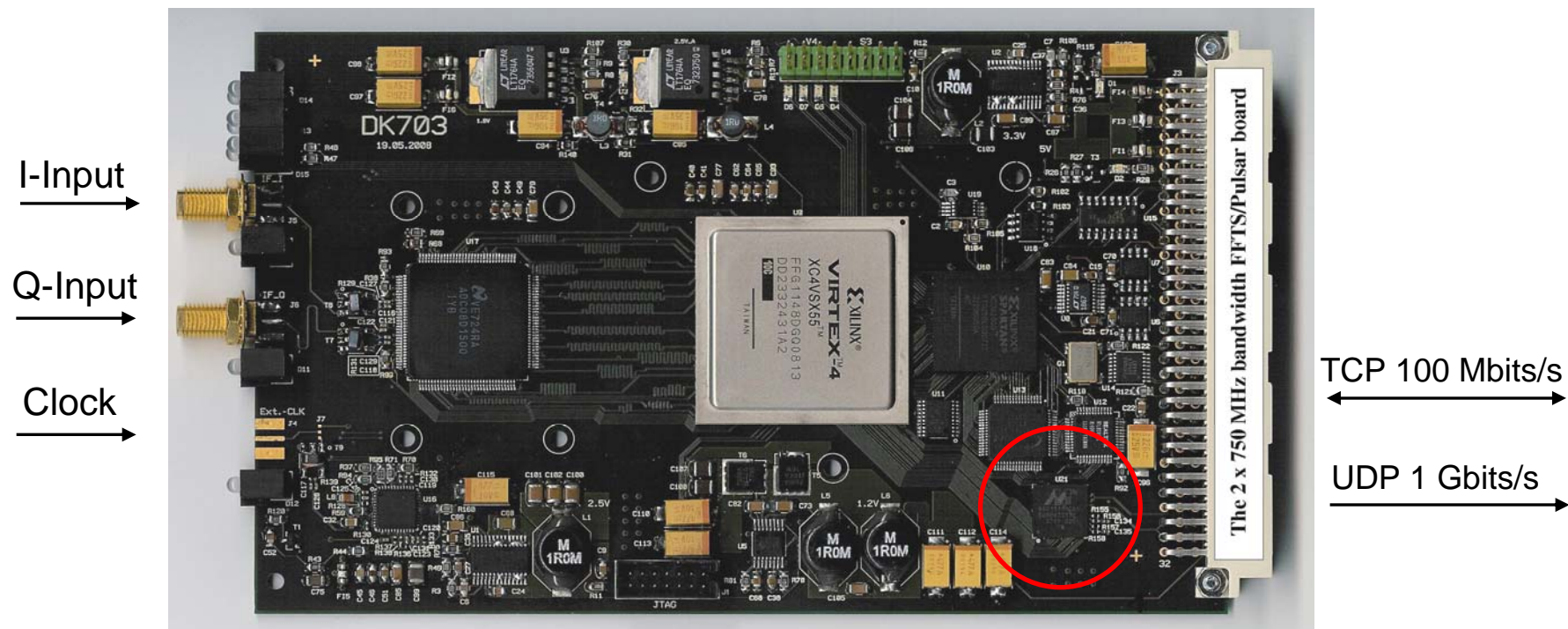
- 2 x 400 MHz bandwidth
- 2 x 65536 frequency channels
- 14-bit DAC dynamic
- arbitrary-function is programmable by Ethernet

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MKIDs Readout: *Prototype Hardware*

ADC-/FPGA-Board – The FFT analyser



Marvell GigaBit chip

- 2 x 750 MHz bandwidth
- 8-bit ADC dynamic
- polyphase FFT processing
- up to 32768 complex (I/Q) frequency channels
- data transfer by 1000 MBit/s Ethernet



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GREAT-FFTS

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Two operation modes:

- **Broadband** : 2x (4x) 1.8 GHz bandwidth / 2x (4x) 8192 channels, ENBW: 255 kHz
- **Narrowband**: 2x (4x) 750 MHz bandwidth / 2x (4x) 16384 channels, ENBW: 53 kHz



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B.Klein ISSTT 2009

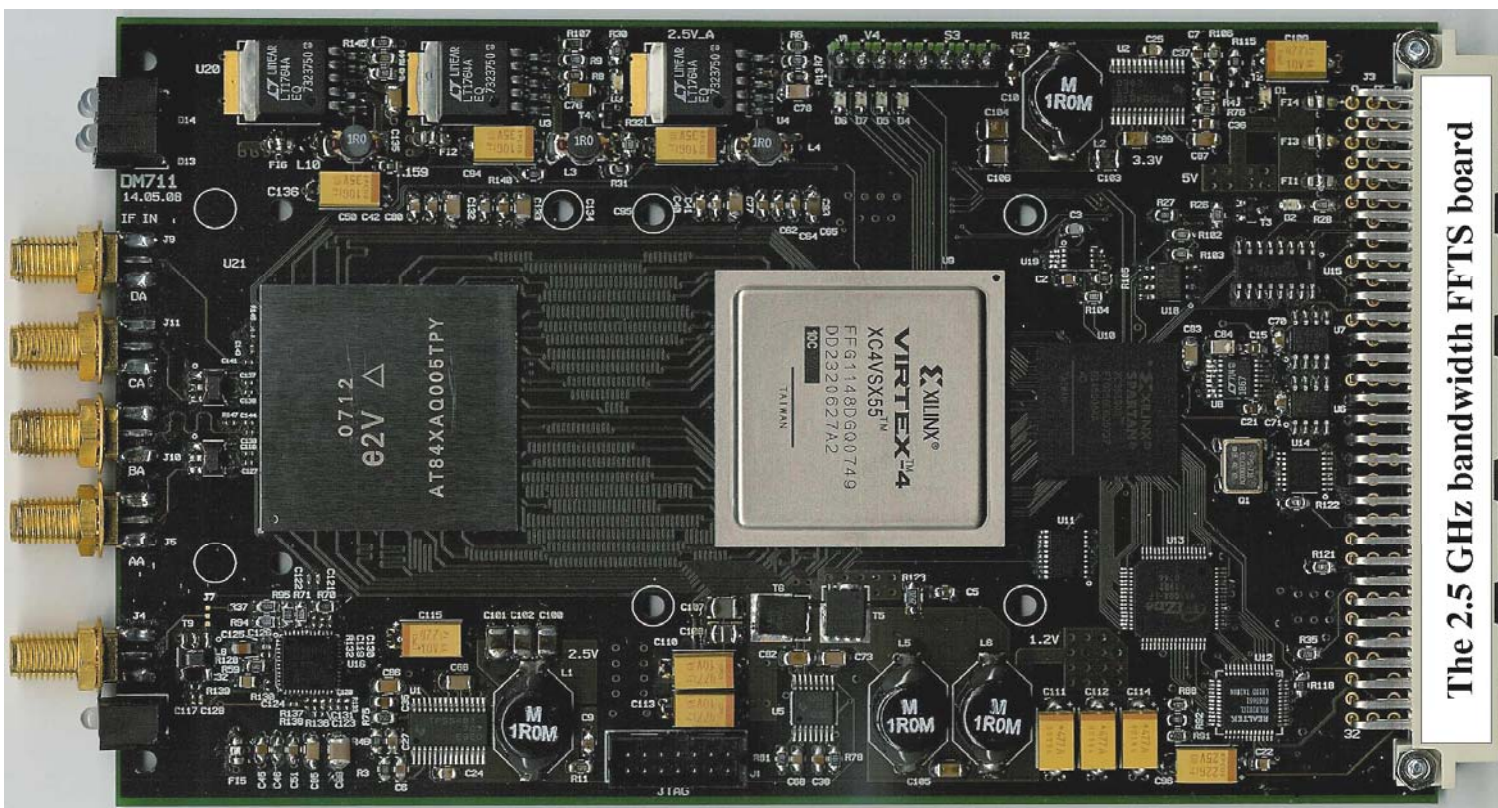


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XFFTS :: The 2.5 GHz test board



XFFTS: 2.5 GHz bandwidth / 256 channels (ENBW 11.3 MHz), 4 IF inputs



E2V 5 GS/s 8-bit ADC, XILINX Virtex-4 SX55



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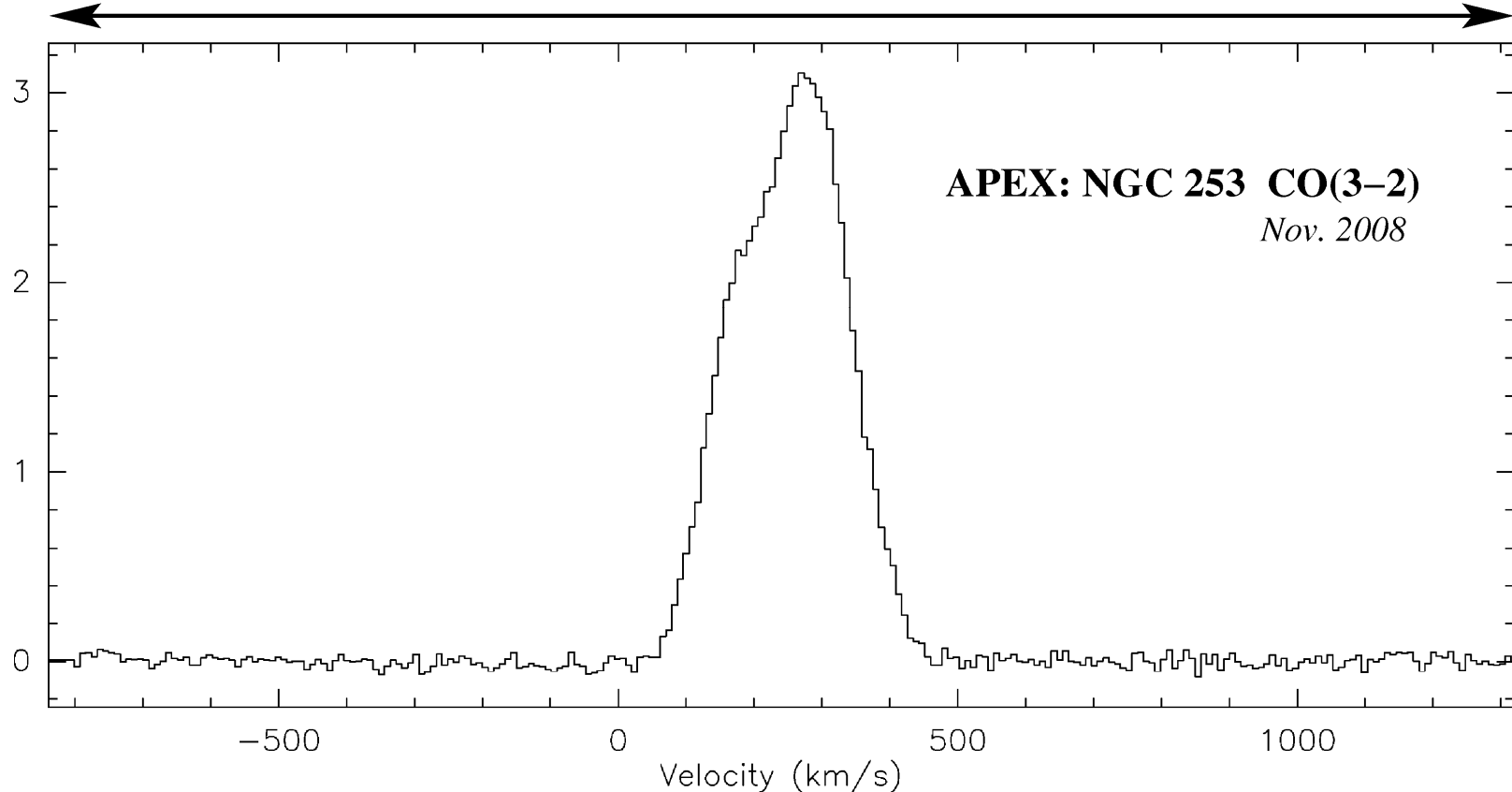


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XFFTS :: *First light at APEX*

2.5 GHz bandwidth



First light spectrum of the new MPIfR 2.5 GHz bandwidth XFFTS towards NGC 253. The CO(3-2) transition at 345 GHz was observed with the APEX-2 facility receiver in November 2008.



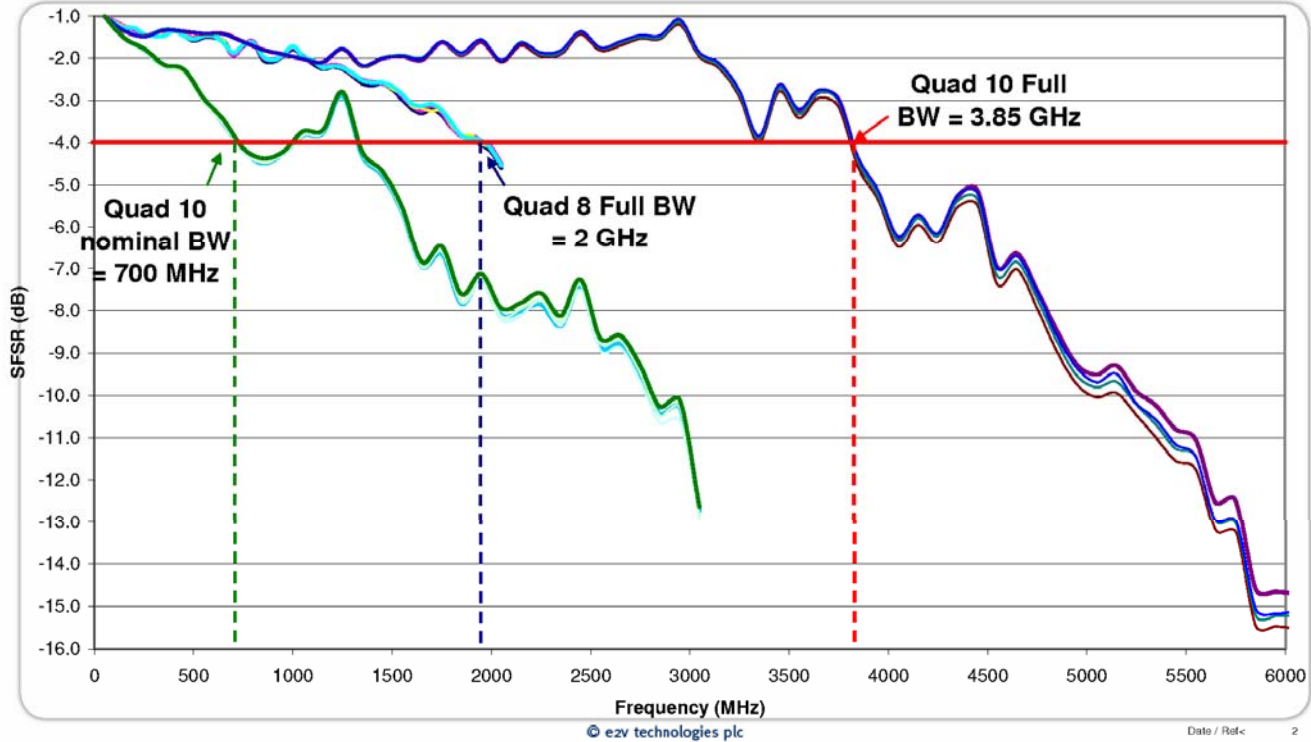
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XFFTS :: The new 10-bit 5 GS/s ADC



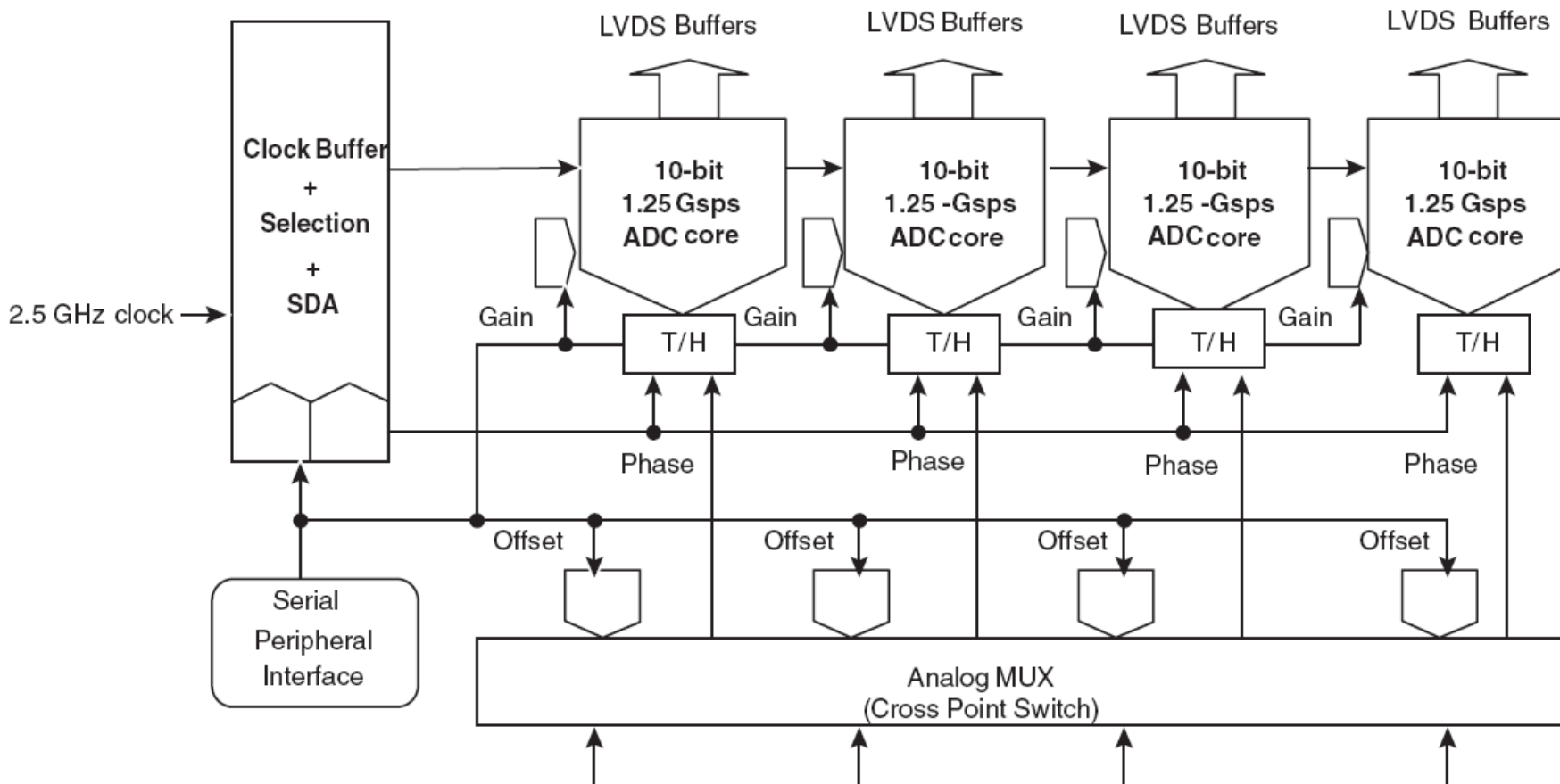
ADC QUAD 10bit **e2v**

Quad 10 Bandwidth @ $F_c = 2.5$ GHz





XFFTS :: The new 10-bit 5 GS/s ADC



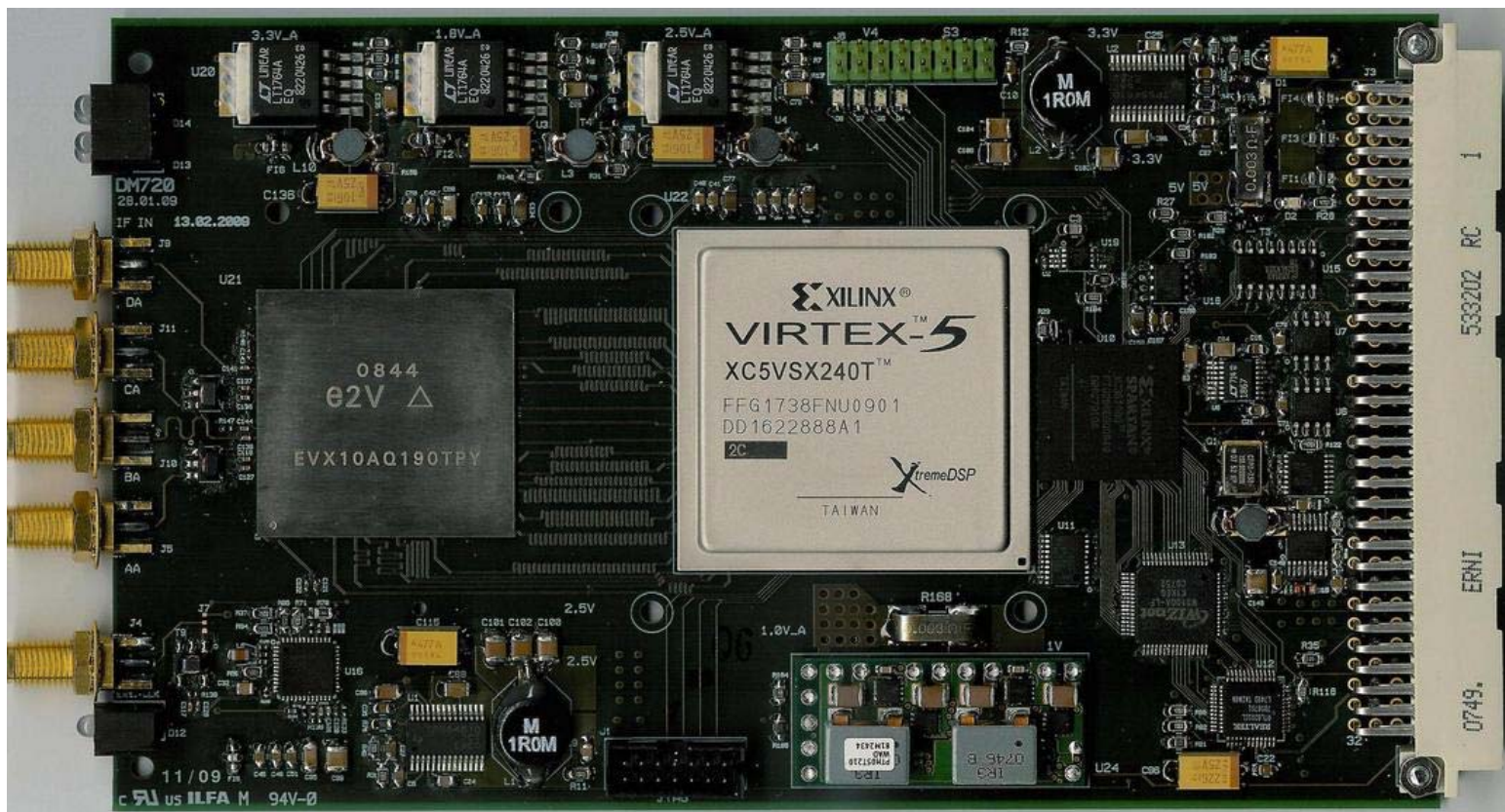


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XFFTS :: *The newest board*



XFFTS: 2.5 GHz bandwidth / 32768 channels (ENBW 88.5 kHz), 4 IF inputs



E2V 5 GS/s 10-bit ADC, XILINX Virtex-5 SX240T



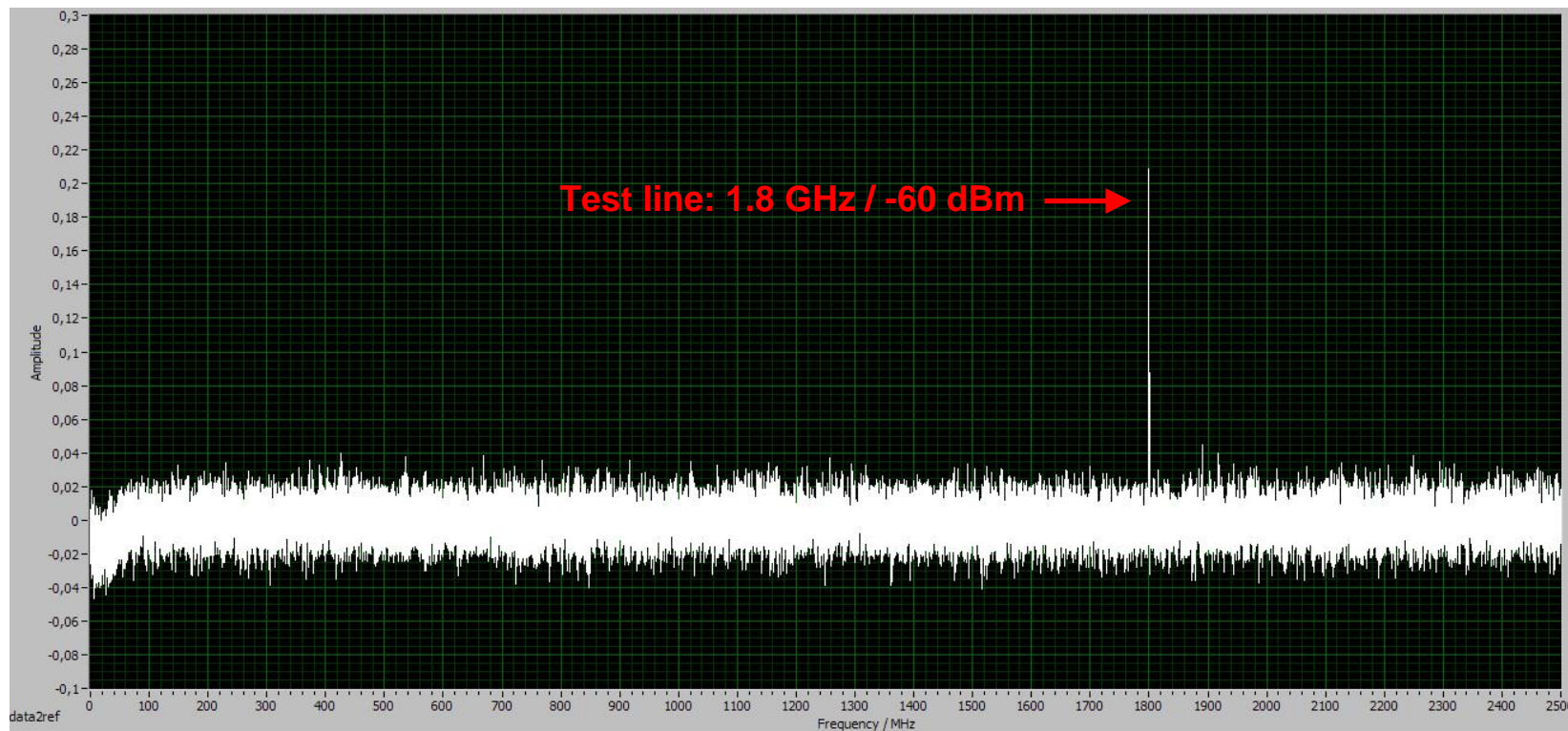
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XFFTS :: *The first lab spectra*

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XFFTS: 2.5 GHz bandwidth / 10-bit / 32768 channels (ENBW 88.5 kHz)



← 2.5 GHz instantaneous bandwidth →
2009/04/22

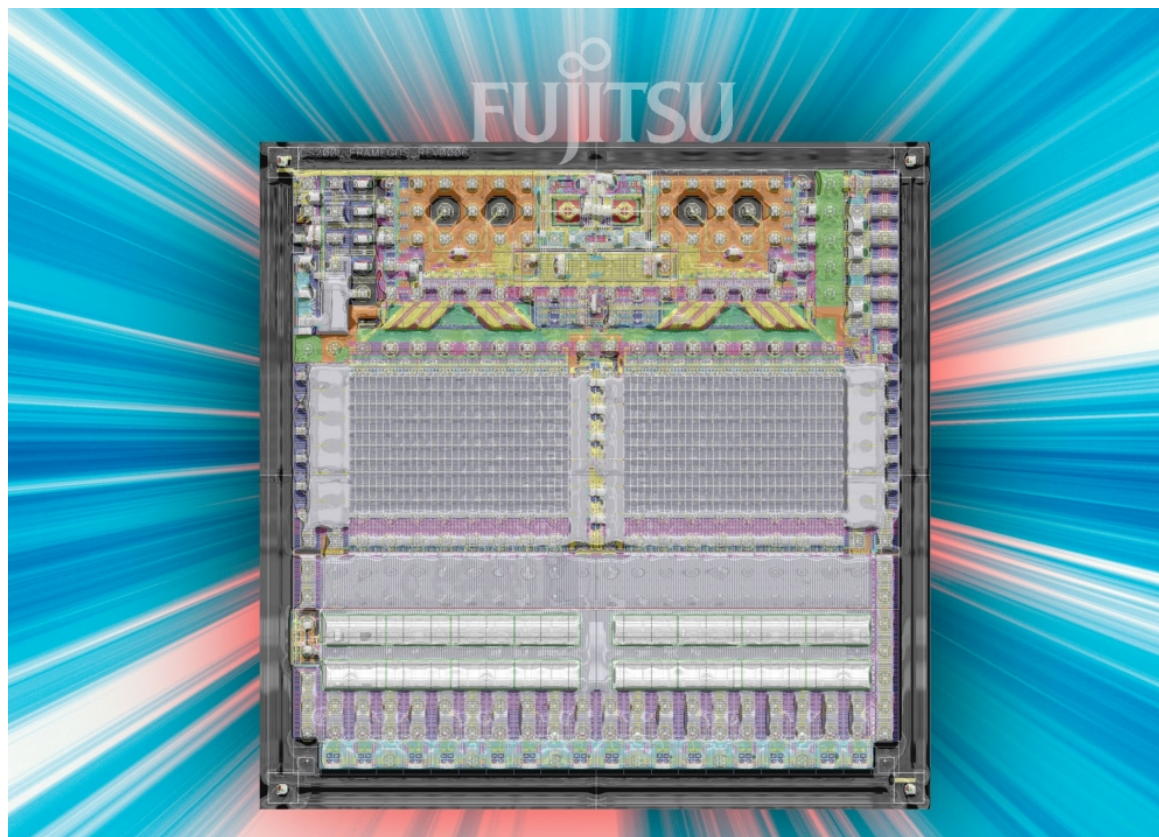
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FFTS :: *An Outlook*

Goal: Sample the whole ALMA IF-bandwidth (4 – 12 GHz)



FUJITSU's 56 GS/s 8-bit ADC, analog input bandwidth (-3 dB): > 15 GHz
(<http://chais.info>)



MPIfR FFTS :: *Summary*

Advantages of our new generation of compact FFT spectrometers:

- ✓ FFTS offer high instantaneous bandwidth up to 2.5 GHz with many thousands frequency channels, thus offering wideband observations with high spectral resolution without the complexity of the IF processing in a hybrid configuration.
- ✓ They provide very high stability by exclusive digital signal processing. Allan stability times of > 1000 seconds have been demonstrated routinely.
- ✓ Our optimized polyphase FFT signal processing pipeline provides a nearly loss-free time to frequency transformation with significant reduced frequency scallop, less noise bandwidth expansion, and faster side lobe fall-off.
- ✓ Field-operations of our FFTS over the last 4 years have proven to be very reliable, with calibration- and aging-free digital processing boards, which are swiftly re-configurable by Ethernet for special observation modes.
- ✓ Low space and power requirements – thus safe to use at high altitude (e.g. APEX at 5100-m) as well as on spacecrafts (Sofia) and future satellites (Millimetron?).
- ✓ Production cost are low compared to traditional spectrometers through use of only commercial components.



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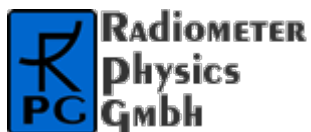
FFTS :: *Contact, Distribution*

Contact:

For further information about the MPIfR FFT spectrometer, future developments and applications, please contact Bernd Klein (bklein@mpifr.de) or Rolf Güsten (rguesten@mpifr.de) at the Max-Planck-Institut für Radioastronomie in Bonn, Germany.



Distribution:



<http://www.radiometer-physics.de>

